

Three Effective Field Theory Vignettes

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Mainz PRISMA Colloquium

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Who Am I?



authors ▾

🔍

astro-ph

hep-ex

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○ 2014-2015
POSTDOC, [Princeton U.](#)

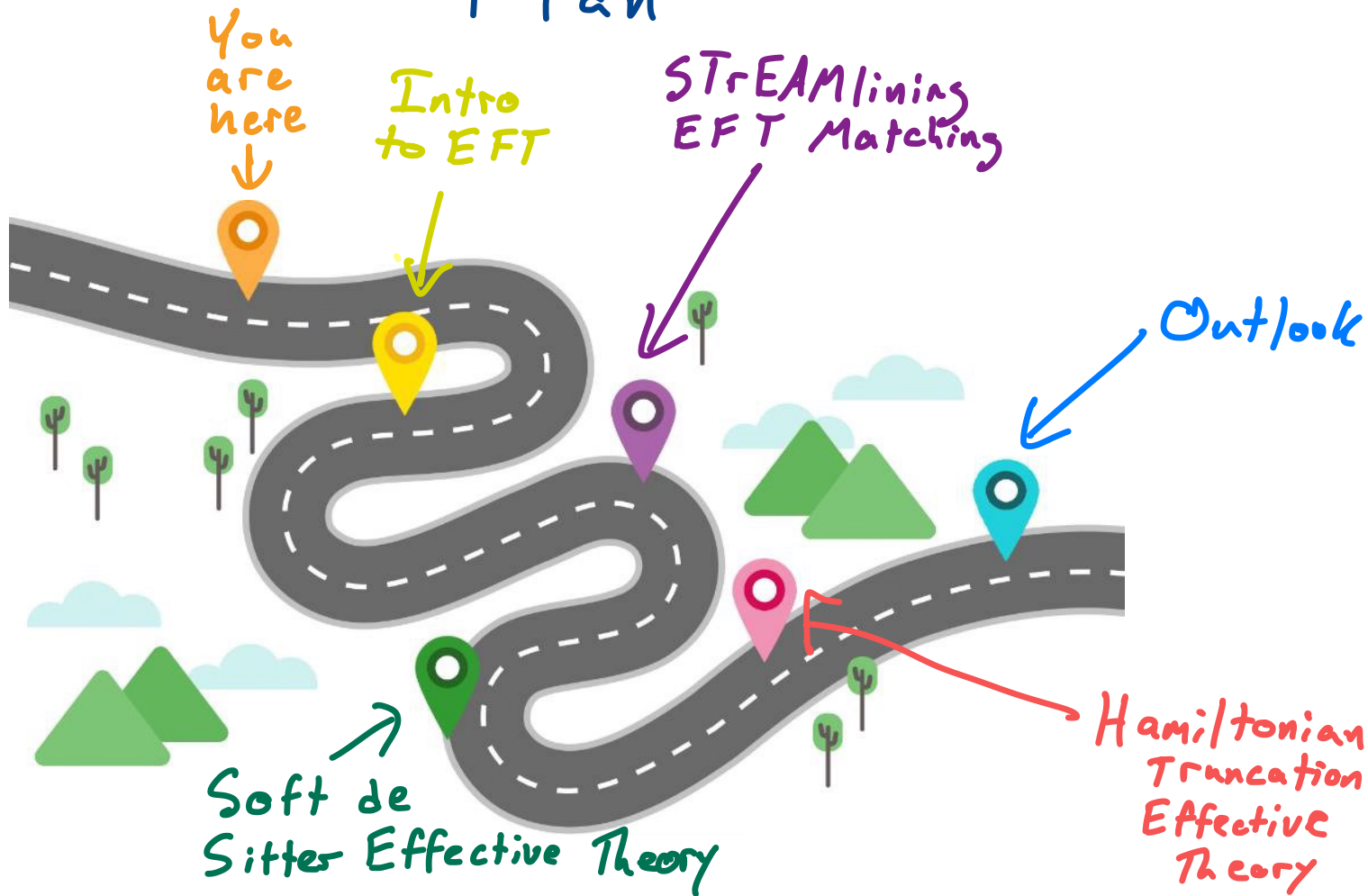
○ 2011-2014
POSTDOC, [SLAC](#)

○ 2006-2011
PHD, [Michigan U.](#)

○ 2002-2006
UNDERGRADUATE, [Alabama U., Huntsville](#)

Effective Field Theory
is everywhere...

Plan



Intro to EFT

As Scales Become Separated: Lectures on Effective Field Theory

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Abstract

These lectures aim to provide a pedagogical introduction to the philosophical underpinnings and technical features of Effective Field Theory (EFT). Improving control of S -matrix elements in the presence of a large hierarchy of physical scales $m \ll M$ is emphasized. Utilizing $\lambda \sim m/M$ as a power counting expansion parameter, we show how matching a UV model onto an EFT makes manifest the notion of separating scales. Renormalization Group (RG) techniques are used to run the EFT couplings from the UV to the IR, thereby resumming large logarithms that would otherwise reduce the efficacy of perturbation theory. A variety of scalar field theory based toy examples are worked out in detail. An approach to consistently evolving a coupling across a heavy particle mass threshold is demonstrated. Applying the same method to the scalar mass term forces us to confront the hierarchy problem. The resummation of a logarithm that lacks explicit dependence on the RG scale is performed. After reviewing the physics of IR divergences, we build a scalar toy version of Soft Collinear Effective Theory (SCET), exposing many subtle aspects of these constructions. We show how SCET can be used to resum the soft and collinear IR Sudakov double logarithms that often appear for processes involving external interacting light-like particles. We conclude with the generalization of SCET to theories of gauge bosons coupled to charged fermions. These lectures were presented at TASI 2018.

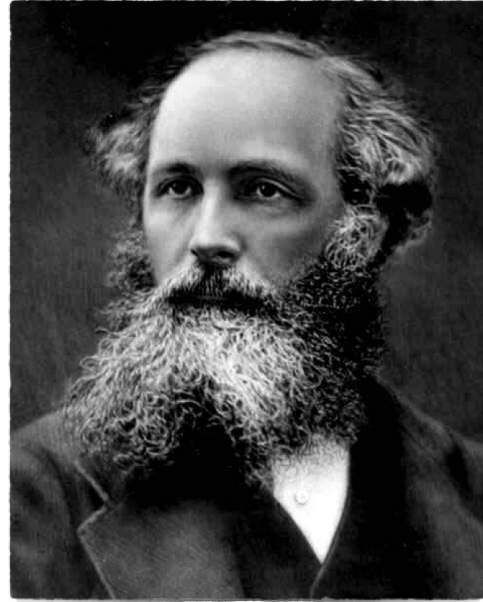
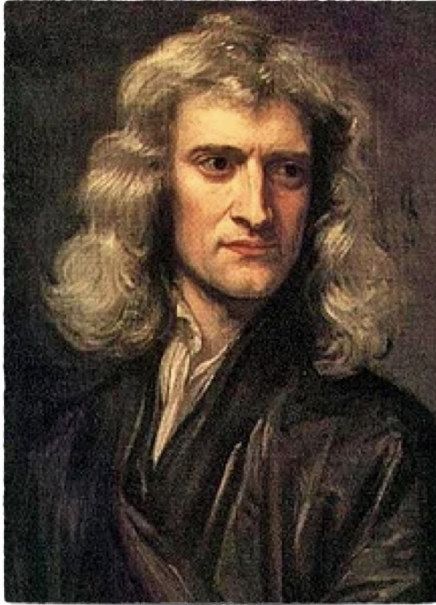
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arXiv:1903.03622

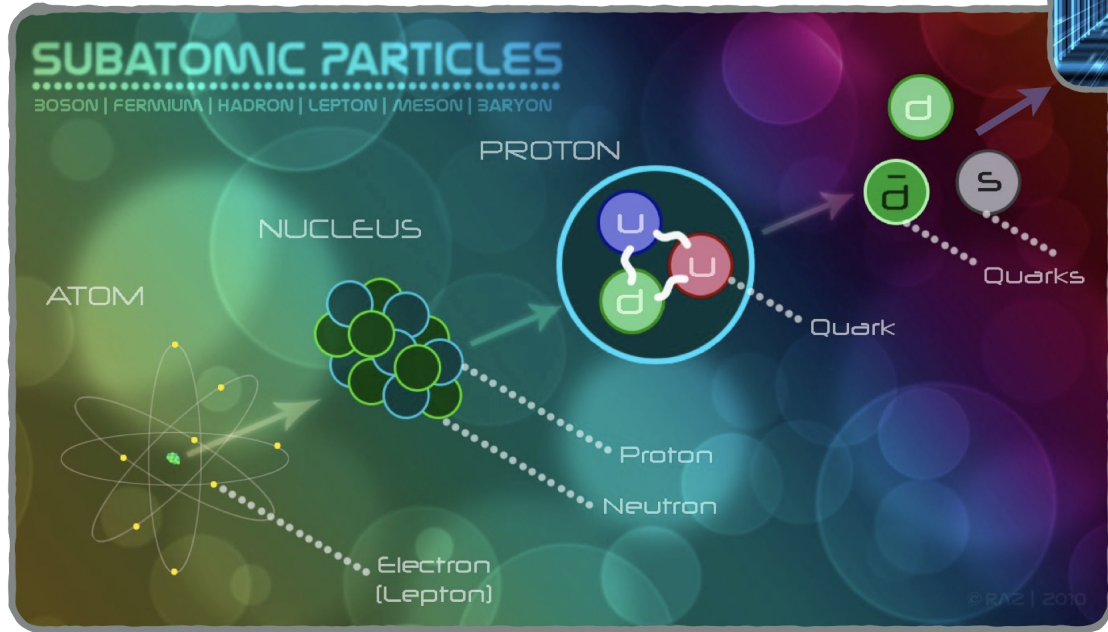
Reductionism

Why didn't Newton and Maxwell need QFT??



"Heavy physics decouples" $\Rightarrow$ Effective description

Reductionism



Large separation of scales

How To Build a Theory

- 1) Degrees of freedom
- 2) Symmetries
- 3) Dimensional analysis



Power Counting

"Physics is essentially dimensional analysis and Taylor expansions"

Large separation of scales

$\Rightarrow$ "power counting parameter" λ

Observables can be computed
order-by-order in power counting

$\Rightarrow$ Predict theoretical uncertainty

Why EFT?

Conceptual: Exposes relevant physics

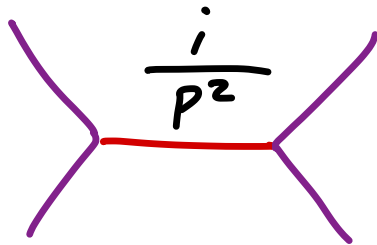
Practical: "Model Independent"
parametrization of low energy physics

Practical: Facilitate hard calculations

Practical: Improve perturbation theory

From High to Low Energy

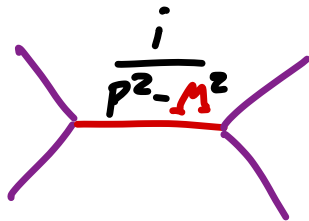
1) $E_{cm} \gg M$



$m \sim 0$
 $M \sim 0$

Theory w/ two massless particles $\Rightarrow$ easy

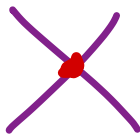
2) $E_{cm} \sim M$



$m \sim 0$
 $M \neq 0$

Multiscale Theory $\Rightarrow$ hard

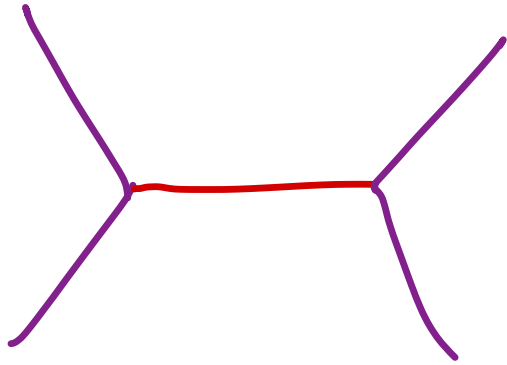
3) $E_{cm} < M$



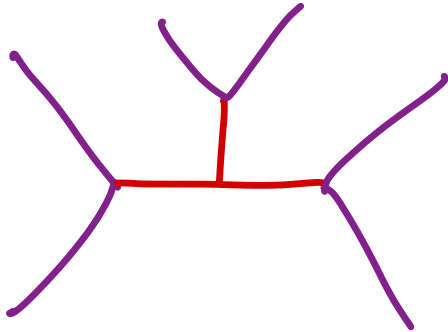
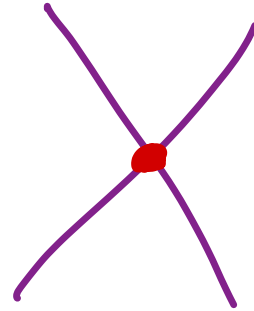
Single particle EFT
 $\Rightarrow$ easy

Heavy Physics Decouples

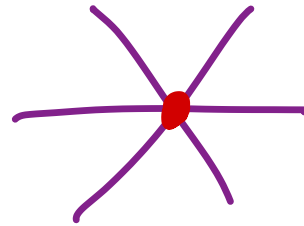
"Integrate out heavy particle"



EFT $\rightarrow$

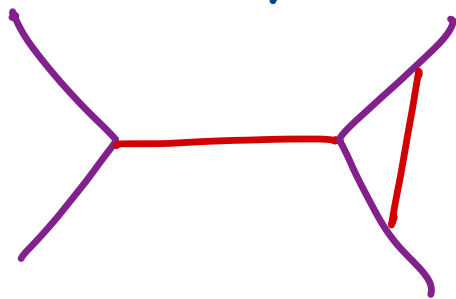


EFT $\rightarrow$



$m \ll M$

Loops



Generate logs e.g. $\log^{m/M}$

Decoupling more subtle

"Matching" full (UV) theory
onto (IR) EFT

EFT and Loops

Loops in QFT $\Rightarrow \log(m/M) \sim \log \lambda$

When $m \ll M$, logs can become large

$\Rightarrow$ must resum them

Promote coupling "constants" to running couplings

Renormalization Group Evolution

EFT

Fundamental (UV) Theory

Matching

M

Running



Renormalization
group evolution

M

Predictions for experiments

Non-relativistic EFT

QFT fields include "particles"
and "anti-particles"

Express $\varphi = \varphi_{\text{particle}} + \varphi_{\text{anti-part}}$

Want observables as expansion in $P/m \ll 1$

$\Rightarrow v \ll 1 \Rightarrow$ power counting

Can "integrate out" $\varphi_{\text{anti-part}} \Rightarrow$ non-rel EFT

Ex: "Heavy Quark Effective Theory"

STrEAMlining EFT Matching

w/ Xiaochuan Lu

+ Zhengkang (Kevin) Zhang

arXiv: 2011.02484

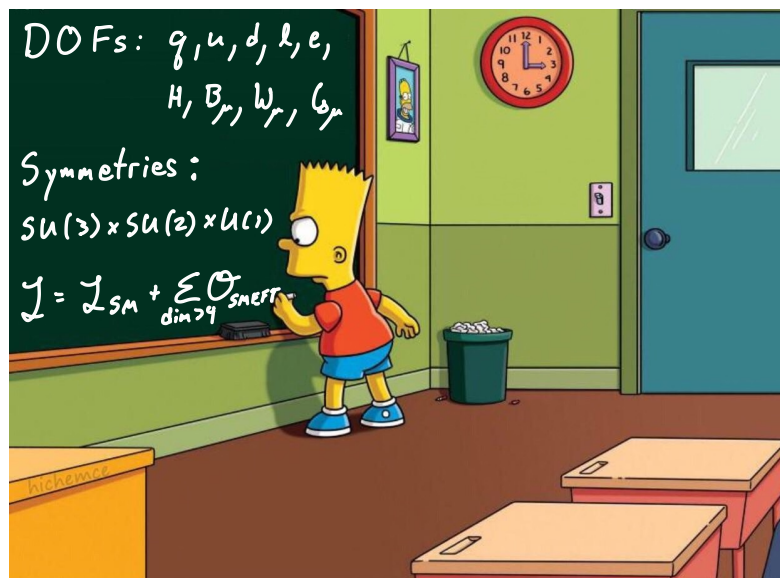
arXiv: 2012.07851

How to organize BSM predictions?

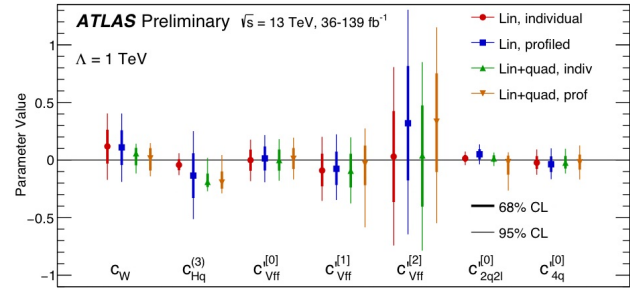
Simplified Models



Effective Field Theory



Constraints on SMEFT



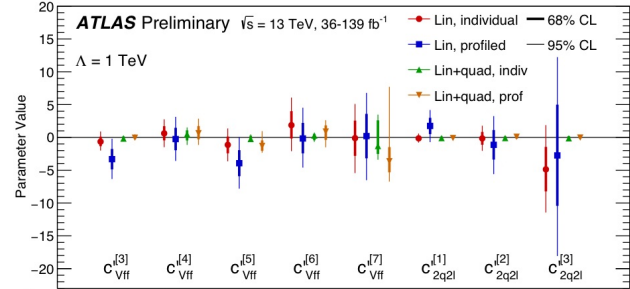
$$C_{Vff}^{[0]} = -0.81c_{HWB}^{(0)} + 0.38c_{HD}^{(0)} + 0.13c_{Ht}^{(0)} + 0.37c_{Hb}^{(0)} - 0.14c_{Hs}^{(0)} + 0.12c_{Hc}^{(0)}$$

$$C_{Vff}^{[1]} = -0.73c_{Ht}^{(1)} - 0.28c_{Hb}^{(1)} - 0.48c_{Hs}^{(1)} + 0.38c_{Hc}^{(1)} + 0.13c_{Hq}^{(1)}$$

$$C_{Vff}^{[2]} = 0.37c_{HWB}^{(2)} + 0.17c_{HD}^{(2)} - 0.31c_{Ht}^{(2)} - 0.53c_{Hb}^{(2)} + 0.25c_{Hs}^{(2)} + 0.59c_{Hc}^{(2)} - 0.21c_{Hq}^{(2)}$$

$$C_{2q2l}^{[0]} = -0.37c_{Ht}^{(0)} + 0.89c_{Hb}^{(0)} - 0.11c_{Hs}^{(0)} - 0.21c_{Hc}^{(0)} - 0.13c_{Hq}^{(0)}$$

$$C_{4q}^{[0]} = 0.11c_{Ht}^{(0)} + 0.22c_{Hb}^{(0)} + 0.95c_{Hs}^{(0)} - 0.2c_{Hc}^{(0)}$$



$$C_{Vff}^{[3]} = -0.19c_{Ht}^{(3)} - 0.14c_{Hb}^{(3)} + 0.86c_{Hs}^{(3)} + 0.41c_{Hc}^{(3)} - 0.17c_{Hq}^{(3)}$$

$$C_{Vff}^{[4]} = -0.35c_{HWB}^{(4)} + 0.49c_{HD}^{(4)} + 0.26c_{Ht}^{(4)} + 0.35c_{Hb}^{(4)} + 0.51c_{Hs}^{(4)} + 0.38c_{Hc}^{(4)} + 0.18c_{Hq}^{(4)}$$

$$C_{Vff}^{[5]} = 0.25c_{HD}^{(5)} + 0.33c_{Ht}^{(5)} - 0.22c_{Hb}^{(5)} + 0.18c_{Hs}^{(5)} - 0.35c_{Hc}^{(5)} - 0.3c_{Hq}^{(5)} + 0.71c_{Hq}^{(5)} - 0.16c_{Hq}^{(5)}$$

$$C_{Vff}^{[6]} = -0.22c_{Ht}^{(6)} + 0.52c_{Hb}^{(6)} - 0.39c_{Hs}^{(6)} + 0.44c_{Hc}^{(6)} - 0.22c_{Hq}^{(6)} + 0.52c_{Hq}^{(6)}$$

$$C_{2q2l}^{[1]} = -0.28c_{HWB}^{(1)} + 0.71c_{HD}^{(1)} - 0.31c_{Ht}^{(1)} - 0.21c_{Hb}^{(1)} - 0.5c_{Hs}^{(1)} - 0.14c_{Hc}^{(1)}$$

$$C_{2q2l}^{[2]} = 0.56c_{Ht}^{(2)} + 0.44c_{Hb}^{(2)} + 0.61c_{Hs}^{(2)} - 0.1c_{Hc}^{(2)} + 0.34c_{Hq}^{(2)}$$

$$C_{2q2l}^{[3]} = 0.68c_{Ht}^{(3)} + 0.15c_{Hb}^{(3)} + 0.33c_{Hs}^{(3)} - 0.51c_{Hc}^{(3)} + 0.13c_{Hq}^{(3)} - 0.37c_{Hq}^{(3)}$$

$$C_{2q2l}^{[4]} = -0.27c_{Ht}^{(4)} + 0.79c_{Hb}^{(4)} - 0.39c_{Hs}^{(4)} + 0.28c_{Hc}^{(4)} - 0.22c_{Hq}^{(4)} - 0.16c_{Hq}^{(4)}$$

SMEFT
 = Standard
 Model
 EFT

How to interpret?

EFT Matching

Connect UV theories to
EFT parameters

Two approaches



Feynman
diagrams

(need EFT in)
advance



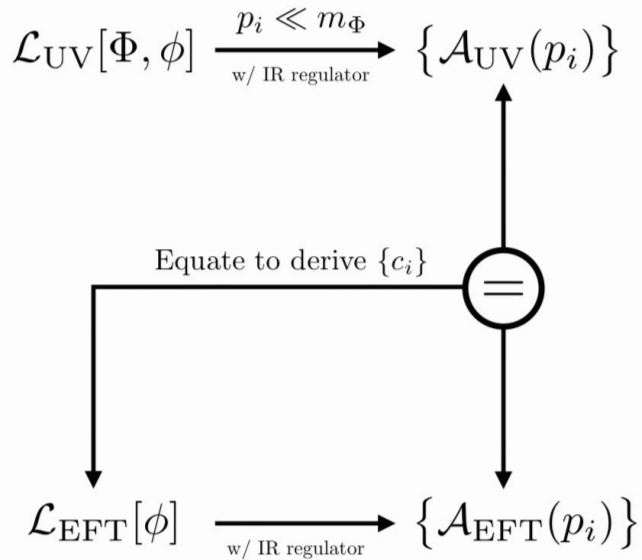
Functional
Methods

(evaluate path
integral directly)

Matching

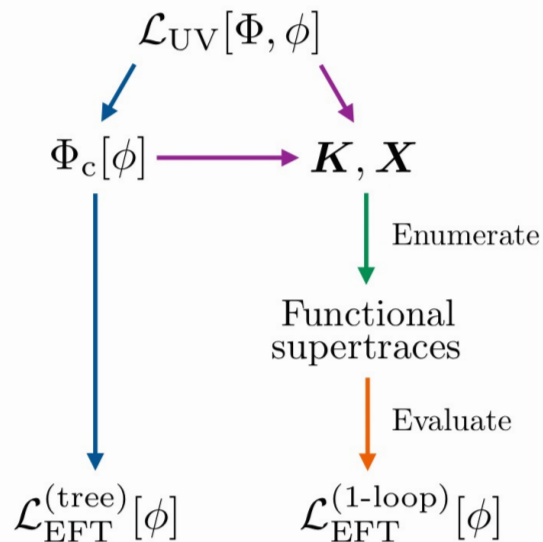
Amplitude matching

(with Feynman diagrams)

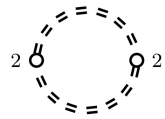


Functional matching

(our prescription)



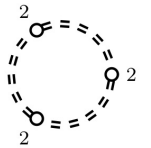
Diagrammatic Prescription



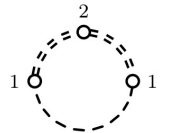
$$= -\frac{i}{2} \frac{1}{2} \text{STr} \left[\left(\frac{1}{P^2 - M^2} U_{SS}^{[2]} \right)^2 \right] \Big|_{\text{hard}},$$



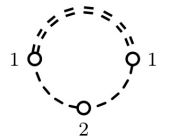
$$= -\frac{i}{2} \text{STr} \left[\frac{1}{P^2 - M^2} U_{SH}^{[1]} \frac{1}{P^2 - m^2} U_{HS}^{[1]} \right] \Big|_{\text{hard}}.$$



$$= -\frac{i}{2} \frac{1}{3} \text{STr} \left[\left(\frac{1}{P^2 - M^2} U_{SS}^{[2]} \right)^3 \right] \Big|_{\text{hard}},$$



$$= -\frac{i}{2} \text{STr} \left[\frac{1}{P^2 - M^2} U_{SS}^{[2]} \frac{1}{P^2 - M^2} U_{SH}^{[1]} \frac{1}{P^2 - m^2} U_{HS}^{[1]} \right] \Big|_{\text{hard}},$$



$$= -\frac{i}{2} \text{STr} \left[\frac{1}{P^2 - M^2} U_{SH}^{[1]} \frac{1}{P^2 - m^2} U_{HH}^{[2]} \frac{1}{P^2 - m^2} U_{HS}^{[1]} \right] \Big|_{\text{hard}}.$$



Super traces
evaluated
using STRIAM
package

Example: Singlet Extended SM

Operator	Coefficient $\times 16\pi^2$
$ H ^2$	$\left[\frac{1}{2}(\kappa M^2 - \mu_S A) + A^2\left(1 + \frac{m^2}{M^2} + \frac{m^4}{M^4}\right)\right]\left(1 - \log \frac{M^2}{\mu^2}\right)$
$ H ^4$	$\frac{\kappa^2}{4}\left(-\log \frac{M^2}{\mu^2}\right) + \frac{\mu_S A}{M^2}\left(\frac{5}{2} - \frac{\mu_S A}{4M^2} + \frac{A^2}{M^2}\right)$ $+ \frac{A^2}{M^2}\left[\left(\frac{\lambda_S}{4} + 3\lambda_H\right)\left(1 - \log \frac{M^2}{\mu^2}\right) - 2\left(\kappa + \frac{A^2}{M^2}\right)\left(\frac{3}{2} - \log \frac{M^2}{\mu^2}\right)\right]$ $+ \frac{m^2}{M^2} \frac{A^2}{M^2}\left[6\lambda_H\left(1 - \log \frac{M^2}{\mu^2}\right) - 3\left(\kappa + \frac{2A^2}{M^2}\right)\left(\frac{4}{3} - \log \frac{M^2}{\mu^2}\right) + \frac{\mu_S A}{M^2}\left(2 - \log \frac{M^2}{\mu^2}\right)\right]$
$ D_\mu H ^2$	$\frac{A^2}{2M^2} + \frac{A^2 m^2}{M^4}\left(\frac{5}{2} - \log \frac{M^2}{\mu^2}\right)$

Operator	Coefficient $\times 16\pi^2$
$ H ^6$	$\frac{1}{M^2}\left(-\frac{\kappa^3}{12} - \frac{\kappa^2 \mu_S A}{4M^2} + \frac{\kappa \mu_S^2 A^2}{2M^4} - \frac{\lambda_S A^4}{2M^4} - \frac{\nu_S^3 A^4}{6M^6} + \frac{\mu_S^3 A^4}{M^6}\right)$ $+ \frac{\kappa A^2}{M^4}\left[3\kappa\left(\frac{11}{6} - \log \frac{M^2}{\mu^2}\right) - \frac{\lambda_S}{6}\left(2 - \log \frac{M^2}{\mu^2}\right)\right]$ $+ \frac{9\lambda_H A^2}{M^4}\left[-\kappa\left(\frac{4}{3} - \log \frac{M^2}{\mu^2}\right) + \lambda_H\left(1 - \log \frac{M^2}{\mu^2}\right)\right]$ $+ \frac{\mu_S A^3}{M^6}\left[-\kappa\left(5 - \log \frac{M^2}{\mu^2}\right) + \frac{\lambda_S}{12}\left(4 - \log \frac{M^2}{\mu^2}\right) + 3\lambda_H\left(2 - \log \frac{M^2}{\mu^2}\right)\right]$ $+ \frac{A^4}{M^6}\left[\frac{2\lambda_S}{3}\left(\frac{31}{12} - \log \frac{M^2}{\mu^2}\right) - 18\lambda_H\left(\frac{4}{3} - \log \frac{M^2}{\mu^2}\right)\right]$ $- \frac{7\mu_S A^5}{2M^8}\left(\frac{15}{4} - \log \frac{M^2}{\mu^2}\right) + \frac{9A^6}{M^8}\left(\frac{43}{27} - \log \frac{M^2}{\mu^2}\right)$
$ H ^2(\partial^2 H ^2)$	$-\frac{\kappa^2}{24M^2} - \frac{5\kappa\mu_S A}{12M^4}$ $+ \frac{A^2}{M^4}\left[2\kappa\left(\frac{17}{12} - \log \frac{M^2}{\mu^2}\right) - \frac{\lambda_S}{2}\left(1 - \log \frac{M^2}{\mu^2}\right) - \frac{\lambda_H}{2}\left(\frac{9}{2} - \log \frac{M^2}{\mu^2}\right)\right]$ $+ \frac{11\mu_S^2 A^2}{24M^6} - \frac{4\mu_S A^3}{3M^6} + \frac{3A^4}{2M^6}\left(\frac{20}{9} - \log \frac{M^2}{\mu^2}\right) - \frac{3\mu_S^2 A^2}{8M^4}\left(\frac{5}{6} - \log \frac{M^2}{\mu^2}\right)$
$ H ^2 D_\mu H ^2$	$\frac{A^2}{M^4}\left[\left(\lambda_H - \frac{A^2}{M^2}\right)\left(\frac{9}{2} - \log \frac{M^2}{\mu^2}\right) - \frac{3\kappa}{2} + \frac{3A^2}{2M^2}\right] - \frac{3\mu_S^2 A^2}{2M^6}\left(\frac{5}{6} - \log \frac{M^2}{\mu^2}\right)$
$\frac{1}{2}(H^\dagger \overleftrightarrow{D}^\mu H)^2$	$\frac{3g_1^2 A^2}{4M^4}\left(\frac{5}{6} - \log \frac{M^2}{\mu^2}\right)$
$ D^2 H ^2$	$\frac{A^2}{6M^4}$

Operator	Coefficient $\times 16\pi^2$
$ig_2(D^\mu H)^\dagger \sigma^I (D^\nu H) W_{\mu\nu}^I$	$-\frac{A^2}{12M^4}$
$ig_1(D^\mu H)^\dagger (D^\nu H) B_{\mu\nu}$	$-\frac{A^2}{12M^4}$
$\frac{ig_2}{2}(H^\dagger \sigma^I \overleftrightarrow{D}^\mu H)(D^\nu W_{\mu\nu}^I)$	$-\frac{A^2}{6M^4}\left(\frac{7}{3} - \log \frac{M^2}{\mu^2}\right)$
$\frac{ig_1}{2}(H^\dagger \overleftrightarrow{D}^\mu H)(\partial^\nu B_{\mu\nu})$	$-\frac{A^2}{6M^4}\left(\frac{7}{3} - \log \frac{M^2}{\mu^2}\right)$
$ H ^2 W_{\mu\nu}^I W^{I\mu\nu}$	$\frac{g_2^2 A^2}{16M^4}$
$ H ^2 B_{\mu\nu} B^{\mu\nu}$	$\frac{g_1^2 A^2}{16M^4}$
$H^\dagger \sigma^I H W_{\mu\nu}^I B^{\mu\nu}$	$\frac{g_1 g_2 A^2}{8M^4}$

Operator	Coefficient $\times 16\pi^2$
$(H^\dagger \sigma^I i \overleftrightarrow{D}_\mu H)(\bar{q} \sigma^I \gamma^\mu q)$	$\frac{A^2}{8M^4}(\mathbf{y}_u \mathbf{y}_u^\dagger + \mathbf{y}_d \mathbf{y}_d^\dagger)\left(\frac{5}{2} - \log \frac{M^2}{\mu^2}\right)$
$(H^\dagger i \overleftrightarrow{D}_\mu H)(\bar{q} \gamma^\mu q)$	$-\frac{A^2}{8M^4}(\mathbf{y}_u \mathbf{y}_u^\dagger - \mathbf{y}_d \mathbf{y}_d^\dagger)\left(\frac{5}{2} - \log \frac{M^2}{\mu^2}\right)$
$(H^\dagger i \overleftrightarrow{D}_\mu H)(\bar{u} \gamma^\mu u)$	$\frac{A^2}{4M^4} \mathbf{y}_b^\dagger \mathbf{y}_b \left(\frac{5}{2} - \log \frac{M^2}{\mu^2}\right)$
$(H^\dagger i \overleftrightarrow{D}_\mu H)(\bar{d} \gamma^\mu d)$	$-\frac{A^2}{4M^4} \mathbf{y}_b^\dagger \mathbf{y}_b \left(\frac{5}{2} - \log \frac{M^2}{\mu^2}\right)$
$(H^\dagger \sigma^I i \overleftrightarrow{D}_\mu H)(\bar{l} \sigma^I \gamma^\mu l)$	$\frac{A^2}{8M^4} \mathbf{y}_e \mathbf{y}_e^\dagger \left(\frac{5}{2} - \log \frac{M^2}{\mu^2}\right)$
$(H^\dagger i \overleftrightarrow{D}_\mu H)(\bar{l} \gamma^\mu l)$	$\frac{A^2}{8M^4} \mathbf{y}_e \mathbf{y}_e^\dagger \left(\frac{5}{2} - \log \frac{M^2}{\mu^2}\right)$
$(H^\dagger i \overleftrightarrow{D}_\mu H)(\bar{e} \gamma^\mu e)$	$-\frac{A^2}{4M^4} \mathbf{y}_e^\dagger \mathbf{y}_e \left(\frac{5}{2} - \log \frac{M^2}{\mu^2}\right)$
$i(\tilde{H}^\dagger (D_\mu H))(\bar{u} \gamma^\mu d) \text{ (+h.c.)}$	$-\frac{A^2}{2M^4} \mathbf{y}_b^\dagger \mathbf{y}_b \left(\frac{5}{2} - \log \frac{M^2}{\mu^2}\right)$
$(H^\dagger \sigma^I H)(\bar{q} \sigma^I i \overleftrightarrow{D} q)$	$-\frac{A^2}{8M^4}(\mathbf{y}_u \mathbf{y}_u^\dagger - \mathbf{y}_d \mathbf{y}_d^\dagger)\left(\frac{1}{2} - \log \frac{M^2}{\mu^2}\right)$
$ H ^2(\bar{q} i \overleftrightarrow{D} q)$	$\frac{A^2}{8M^4}(\mathbf{y}_u \mathbf{y}_u^\dagger + \mathbf{y}_d \mathbf{y}_d^\dagger)\left(\frac{1}{2} - \log \frac{M^2}{\mu^2}\right)$
$ H ^2(\bar{u} i \overleftrightarrow{D} u)$	$\frac{A^2}{4M^4} \mathbf{y}_b^\dagger \mathbf{y}_b \left(\frac{1}{2} - \log \frac{M^2}{\mu^2}\right)$
$ H ^2(\bar{d} i \overleftrightarrow{D} d)$	$\frac{A^2}{4M^4} \mathbf{y}_d^\dagger \mathbf{y}_d \left(\frac{1}{2} - \log \frac{M^2}{\mu^2}\right)$
$(H^\dagger \sigma^I H)(\bar{l} \sigma^I i \overleftrightarrow{D} l)$	$\frac{A^2}{8M^4} \mathbf{y}_e \mathbf{y}_e^\dagger \left(\frac{1}{2} - \log \frac{M^2}{\mu^2}\right)$
$ H ^2(\bar{l} i \overleftrightarrow{D} l)$	$\frac{A^2}{8M^4} \mathbf{y}_e \mathbf{y}_e^\dagger \left(\frac{1}{2} - \log \frac{M^2}{\mu^2}\right)$
$ H ^2(\bar{e} i \overleftrightarrow{D} e)$	$\frac{A^2}{4M^4} \mathbf{y}_e^\dagger \mathbf{y}_e \left(\frac{1}{2} - \log \frac{M^2}{\mu^2}\right)$
$ H ^2 \bar{q} u \tilde{H} \text{ (+h.c.)}$	$\frac{A^2}{M^4} \mathbf{y}_u \mathbf{y}_b^\dagger \mathbf{y}_b \left(1 - \log \frac{M^2}{\mu^2}\right)$
$ H ^2 \bar{q} d \tilde{H} \text{ (+h.c.)}$	$\frac{A^2}{M^4} \mathbf{y}_d \mathbf{y}_b^\dagger \mathbf{y}_b \left(1 - \log \frac{M^2}{\mu^2}\right)$
$ H ^2 \bar{l} e \tilde{H} \text{ (+h.c.)}$	$\frac{A^2}{M^4} \mathbf{y}_e \mathbf{y}_e^\dagger \mathbf{y}_e \left(1 - \log \frac{M^2}{\mu^2}\right)$

One loop matching is "solved"

Stay Tuned

Functional matching relies
on dim reg and method of regions.

What about γ_5 ???

We have developed a novel
4D regulator for the anomaly

Facilitates integrating out
Weyl fermions.

Soft de Sitter Effective Theory

w/ Dan Green (UCSD)

arXiv: 2007.03693

+ Akhil Premkumar + Alec Ridgway

arXiv: 2106.09728, 2111.09332

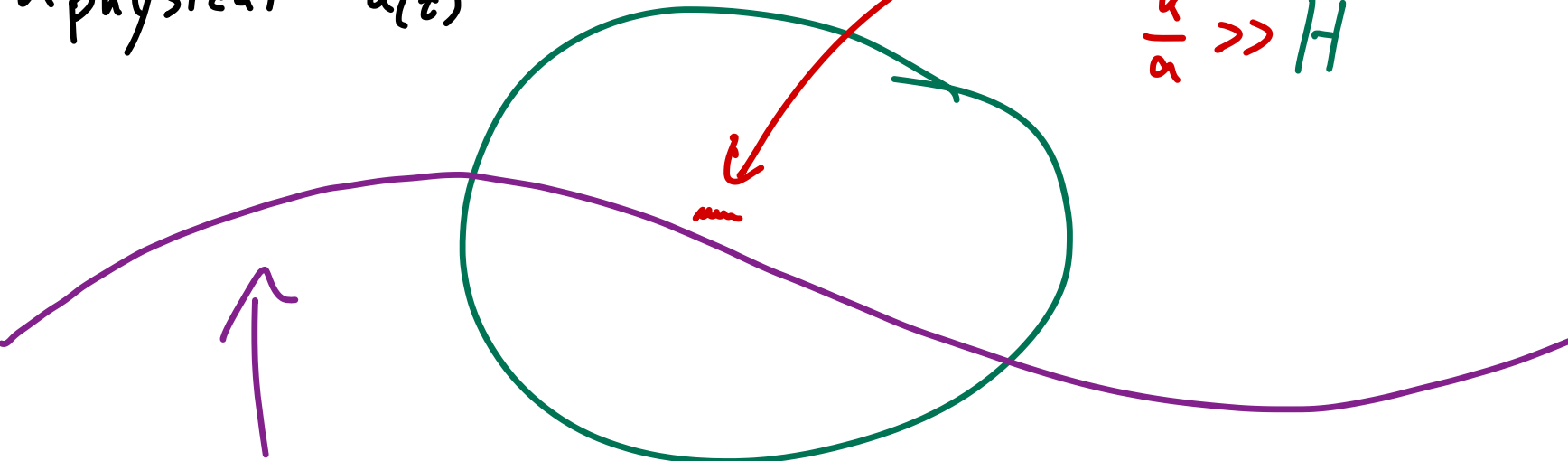
+ new paper today!

Soft de Sitter Effective Theory

$$k_{\text{physical}} = \frac{k}{a(t)}$$

UV modes

$$\frac{k}{a} \gg H$$



SDSET

$$\frac{k}{a} \ll H$$

$$\tau H^{-1}$$

Power counting

$$\lambda = \frac{k}{aH}$$

Scalar Fields in dS

EOM $\ddot{\phi} + 3\dot{\phi} + \frac{k^2}{(aH)^2} \phi + \frac{m^2}{H^2} \phi = 0$

Soft limit $\phi_S = (aH)^{-3/2+\nu} \varphi_S$

w/ $\nu = \pm \sqrt{\frac{9}{4} - \frac{m^2}{H^2}}$

or $\alpha = \frac{3}{2} - \nu$ $\beta = \frac{3}{2} + \nu$ s.t. $\alpha + \beta = 3$

$m \rightarrow 0 \Rightarrow \alpha \rightarrow 0$

One-to-many Mode Expansion

Factorize into soft and hard modes

$$\phi(\vec{x}, t) = \phi_S(\vec{x}, t) + \Phi_H(\vec{x}, t)$$

Integrate out hard modes

$\Rightarrow$ Local operator expansion

Observables order-by-order in power counting

SDSET Fields

Two IR degrees of freedom

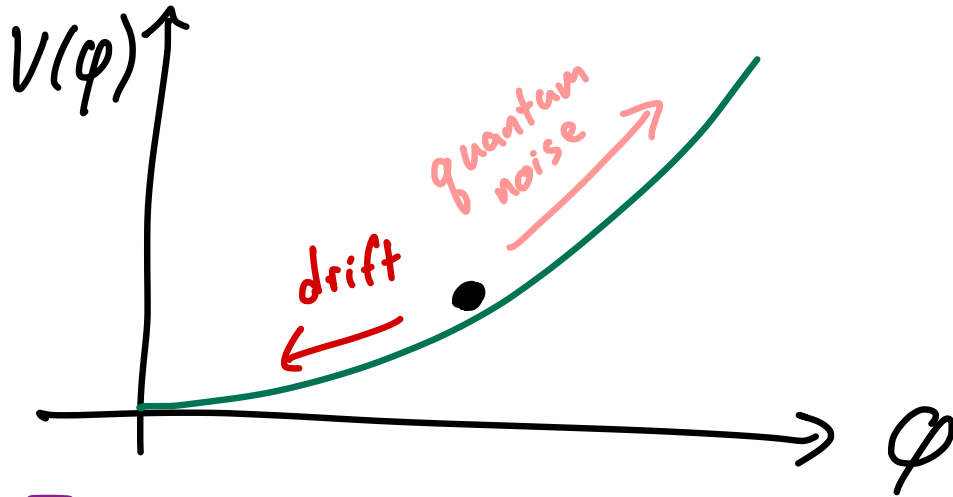
- "Growing" mode φ_+ $\leftarrow$ Correlators of interest
- "Decaying" mode φ_-

w/ $\phi_s = H \left((aH)^{-\alpha} \varphi_+ + (aH)^{-\beta} \varphi_- \right)$

Time dependence factorizes $\therefore$

Starobinsky's Stochastic Inflation

Massless scalar field in dS (1986)

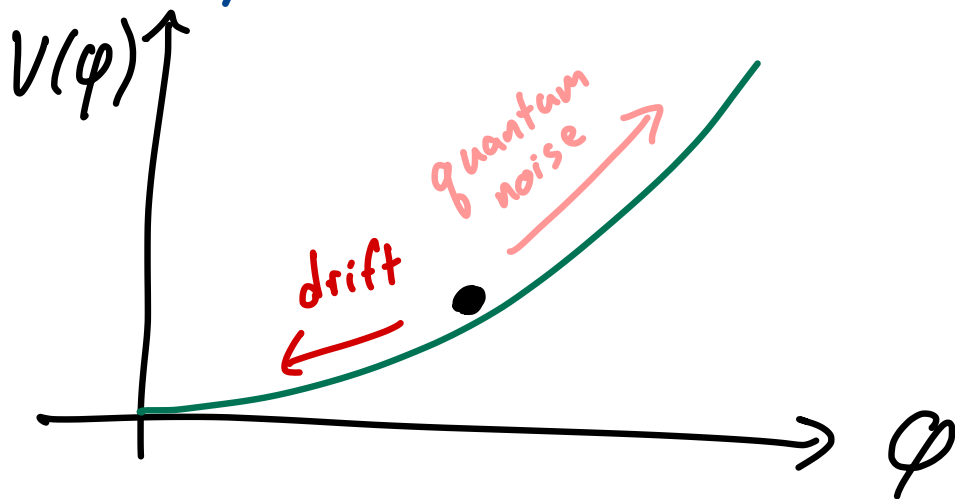


$\left(\begin{array}{l} \phi \text{ is} \\ \text{not the} \\ \text{inflaton} \end{array} \right)$

$\Rightarrow$ Fokker - Planck equation:

$$\frac{\partial}{\partial t} P(\phi, t) = \frac{H^3}{8\pi^2} \frac{\partial^2}{\partial \phi^2} P(\phi, t) + \frac{1}{3H} \frac{\partial}{\partial \phi} [V'(\phi) P(\phi, t)]$$

Starobinsky's Stochastic Inflation

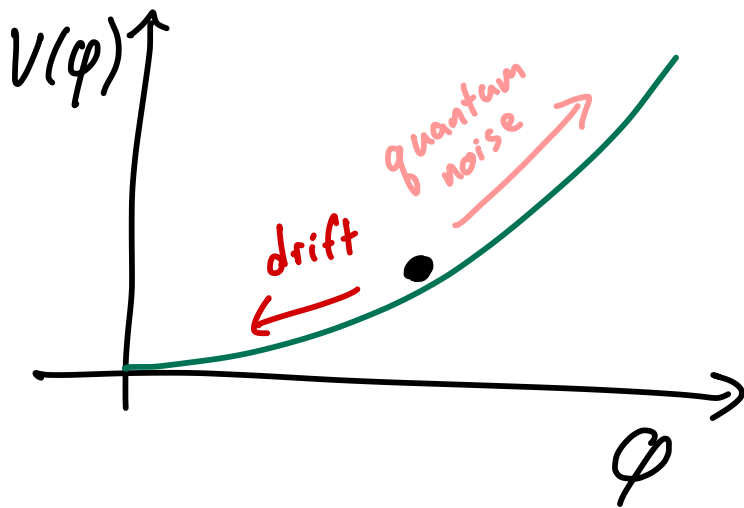


Gaussian noise

Tree-level potential

Systematic corrections?

Stochastic Inflation $\Leftrightarrow$ RG flow



$\Leftrightarrow$

$a_H \frac{\text{UV Theory}}{\text{SdSET}}$

$\downarrow$ RG

Stochastic
Inflation

Light Scalars in dS

Composite operators $n \rightarrow 0$
 $\Rightarrow \alpha \rightarrow 0$

$$\mathcal{O}_n = \phi^n \sim (k/aH)^{n\alpha} \rightarrow \mathcal{O}(1)$$

RG mixing expected

Contract any two legs

$$\langle \mathcal{O}_n \dots \rangle \supset \langle \mathcal{O}_{n-2} \dots \rangle \binom{n}{2} \frac{C_\alpha^2}{2} \int \frac{d^3 p}{(2\pi)^3} \frac{H^{2-2\alpha}}{p^{3-2\alpha}}$$

Light Scalars in dS

$$\int \frac{d^3 p}{(2\pi)^3} \frac{H^{2-2\alpha}}{p^{3-2\alpha}} \quad \text{is scaleless and diverges as } \alpha \rightarrow 0$$

Isolate UV divergence

$$p^2 \rightarrow p^2 + \overline{k}_{\text{IR}}^2$$

$$\langle \mathcal{O}_n \dots \rangle \supset \langle \mathcal{O}_{n-2} \dots \rangle \binom{n}{2} \frac{C_\alpha^2}{4\pi^2} \left(-\frac{1}{2\alpha} - \gamma_E - \log \frac{aH}{\overline{k}_{\text{IR}}} \right)$$

Dynamical RG $\Leftrightarrow$ Stochastic Inflation

Resum time dependent logs:

$$\frac{\partial}{\partial t} \langle \phi_n \dots \rangle = -\frac{n}{3} \sum_{m \geq 1} \frac{c_m}{m!} \langle \phi_{n-1} \phi_m \dots \rangle$$

insertion
of potential
 $V \sim c_m \phi_+^m \phi_-$

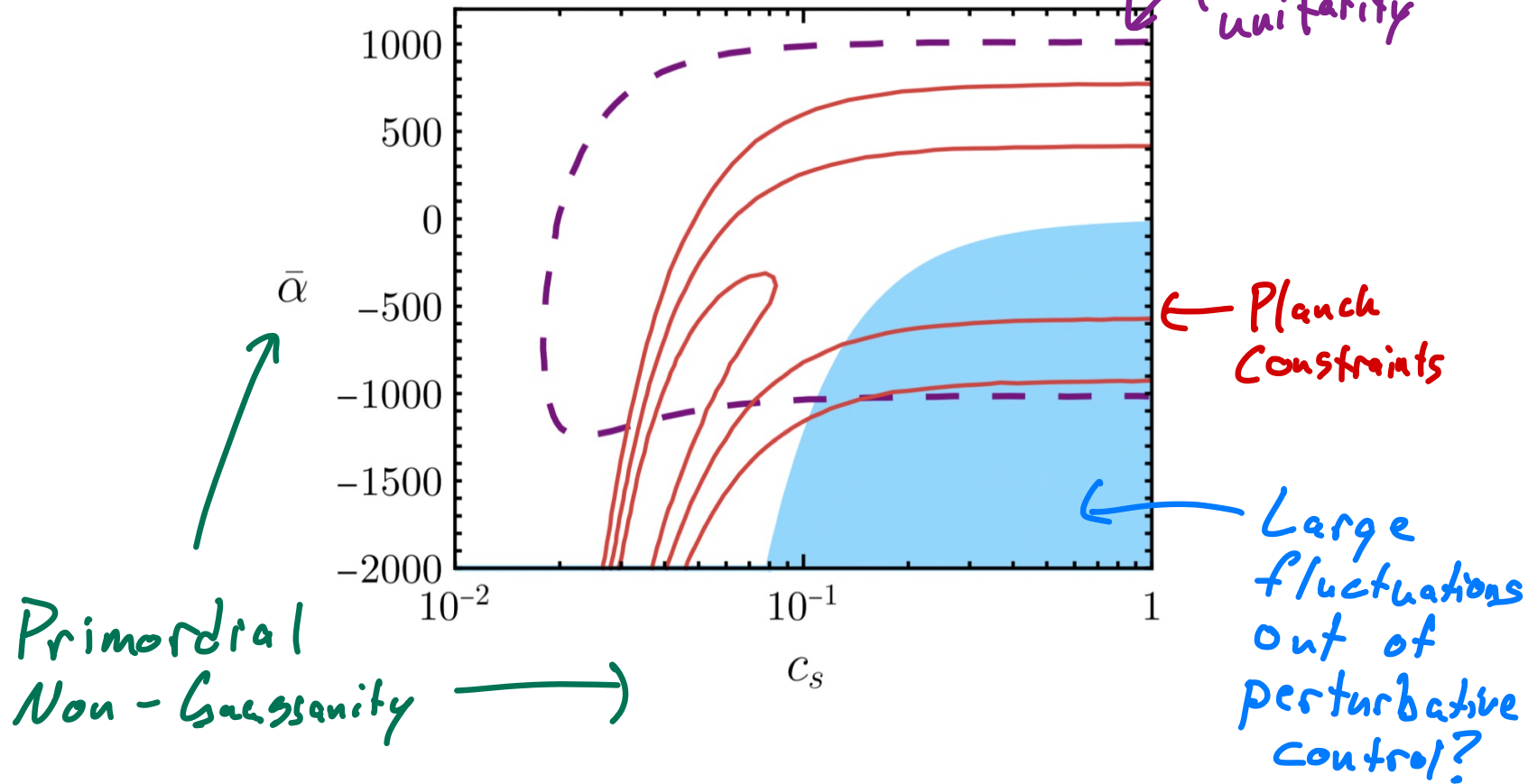
$$+ \frac{n(n-1)}{8\pi^2} \langle \phi_{n-2} \dots \rangle$$

(Starobinsky; Starobinsky, Yokoyama)

Is equivalent to a Fokker-Planck eq

for $P(\phi, t)$ w/ $\langle \phi^n \rangle = \int d\phi P(\phi, t) \phi^n$ (Baumgart + Sundrum)

Apply Same Techniques to Inflaton



Large Deviations in the Early Universe

We have developed a Out today!

comprehensive understanding for the origin of this breakdown.

The tails of these distributions are dominated by a different saddle point $\Rightarrow$ operator scalings change and UV sensitivity emerges.

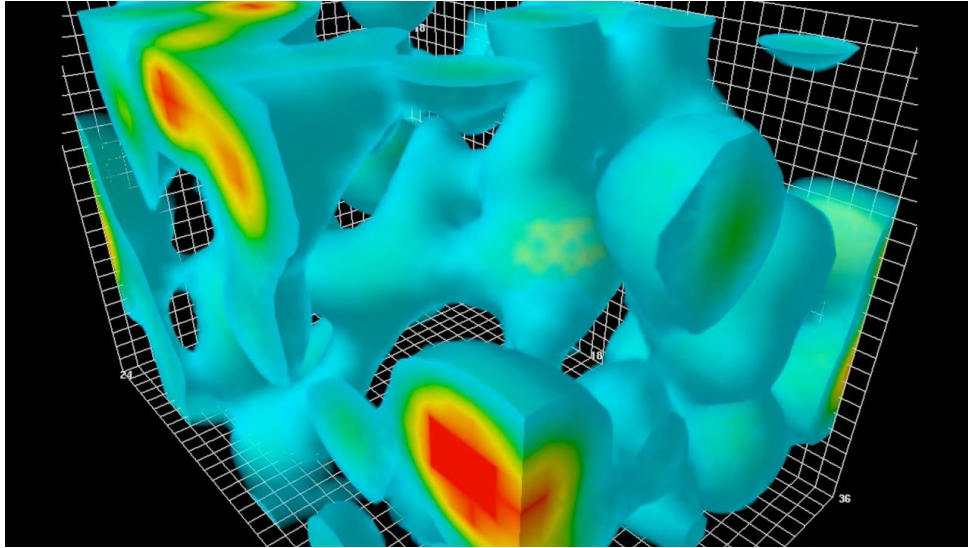
Hamiltonian Truncation Effective Theory

w/ Kara Farnsworth
Rachel Houtz
Markus Luty

arXiv: 2110.08273

Explore Strongly Coupled QFT

Lattice QCD



(from Adelaide group)

Hamiltonian Truncation

1) Write $\hat{H} = \hat{H}_0 + \hat{V}$ where H_0
can be solved exactly

$$\hat{H}_0 |E_i\rangle = E_i |E_i\rangle$$

2) Introduce "energy cutoff" $E_{\max}$

3) Compute matrix elements of $\hat{H}$
using truncated basis

4) Diagonalize $\Rightarrow$ approx energy spectrum

Want to Study IR Strong Theories

Simple case study

$\lambda \phi^4$ Theory in 2D

Dimensional analysis

$$[\phi] = 0 \Rightarrow [\lambda] = 2$$

Relevant operator

$\Rightarrow$ weak in UV and strong in IR

Improve Numerical Predictions

Interpret $E_{\max}$ as

EFT cutoff scale

Power counting $\lambda \sim E_{\text{IR}}/E_{\max}$

Write $H_{\text{eff}} = H_0 + H_1 + H_2 + \dots$

w/ $H_n = O(V^n)$ finite volume non-locality

Expect $H_2 \sim \frac{\lambda^2}{E_{\max}^2} \int dx \left[\phi^2 + \phi^4 + \frac{R^{-1} + H_0}{E_{\max}} (1 + \phi^2 + \phi^4) \right] + \dots$

EFT Expectations

$$H_2 \sim \frac{\lambda^2}{E_{\max}^2} \int dx \left[\underbrace{\phi^2 + \phi^4}_{\text{"local approx"}} + \frac{R^{-1} + H_0}{E_{\max}} (1 + \phi^2 + \phi^4) + \dots \right]$$

Compute energy eigenvalues
as a function of $E_{\max}$

"Raw truncation" converges as $1/E_{\max}^2$

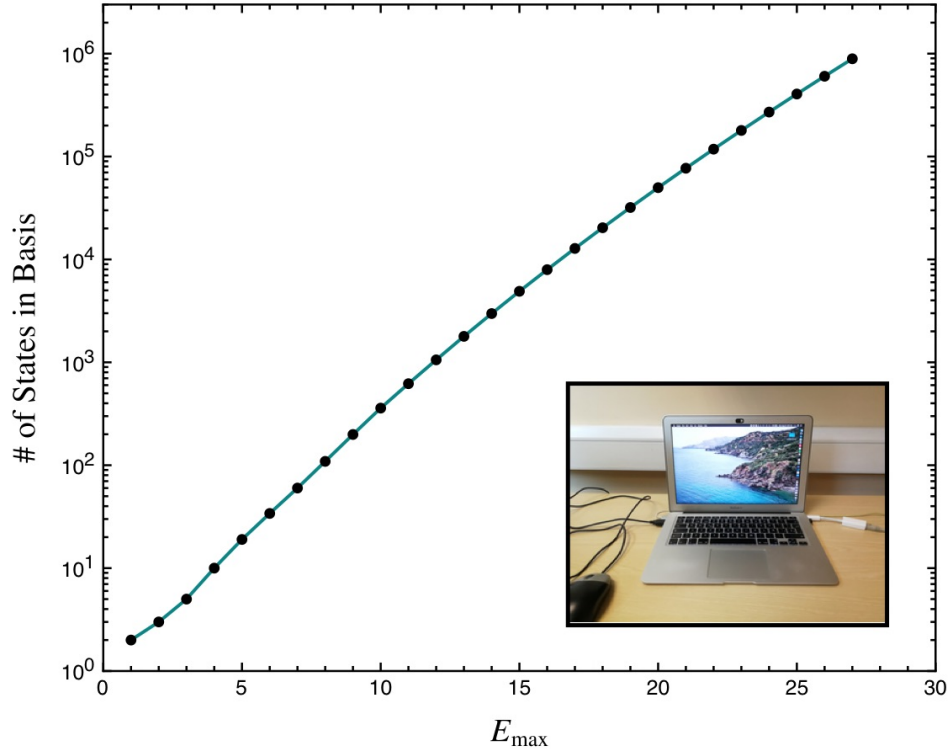
"Improved truncation" converges as $1/E_{\max}^3$

Technical Details

- Need IR cutoff $\Rightarrow$ work in finite volume
- Observable for matching is "transition matrix" (old fashioned perturbation theory)
- Derived set of diagrammatic rules
- $E_{\max}$ breaks Lorentz invariance
 - $\Rightarrow$ Sensitive to states that do not participate in interaction.

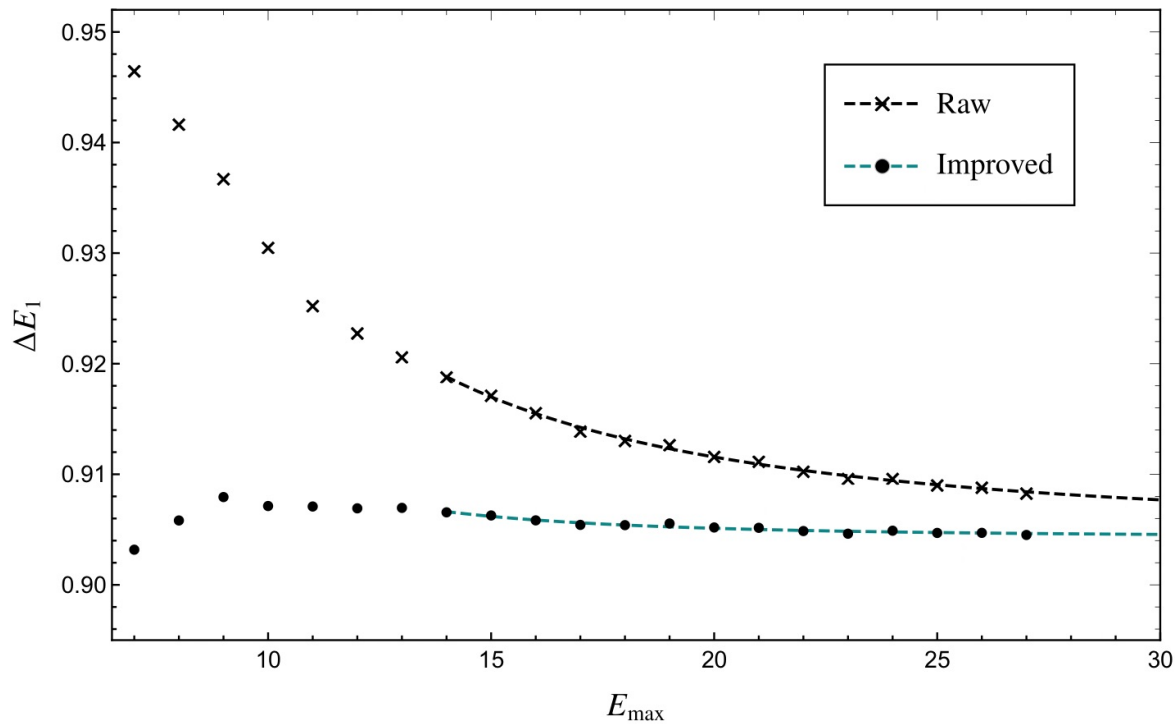
Numerical Results

$$m_{\text{NO}} = 1, 2\pi R = 10$$

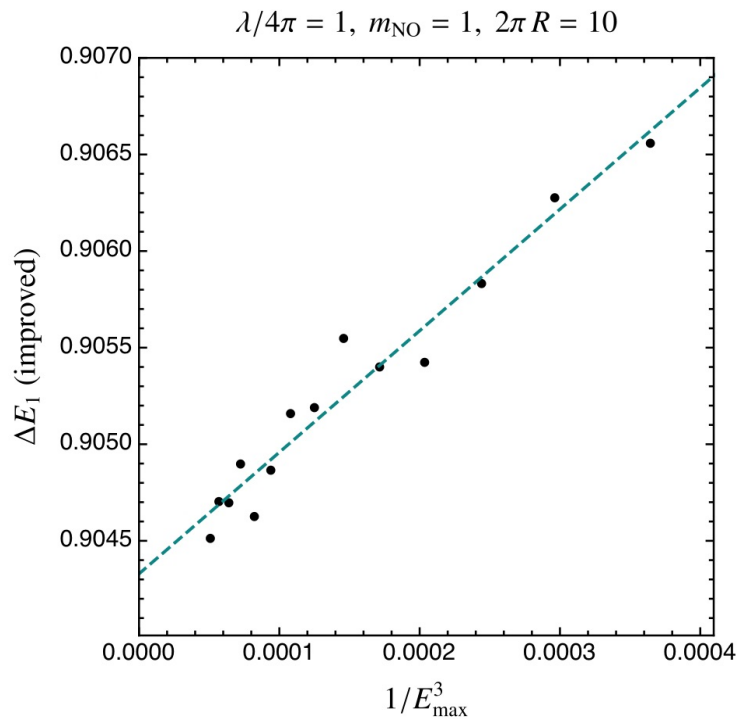
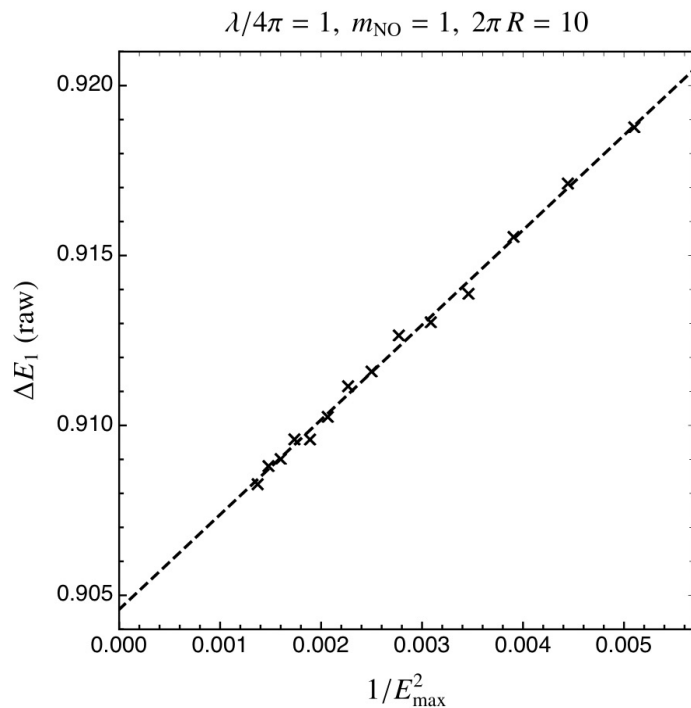


Numerical Results

$$\lambda/4\pi = 1, m_{\text{NO}} = 1, 2\pi R = 10$$

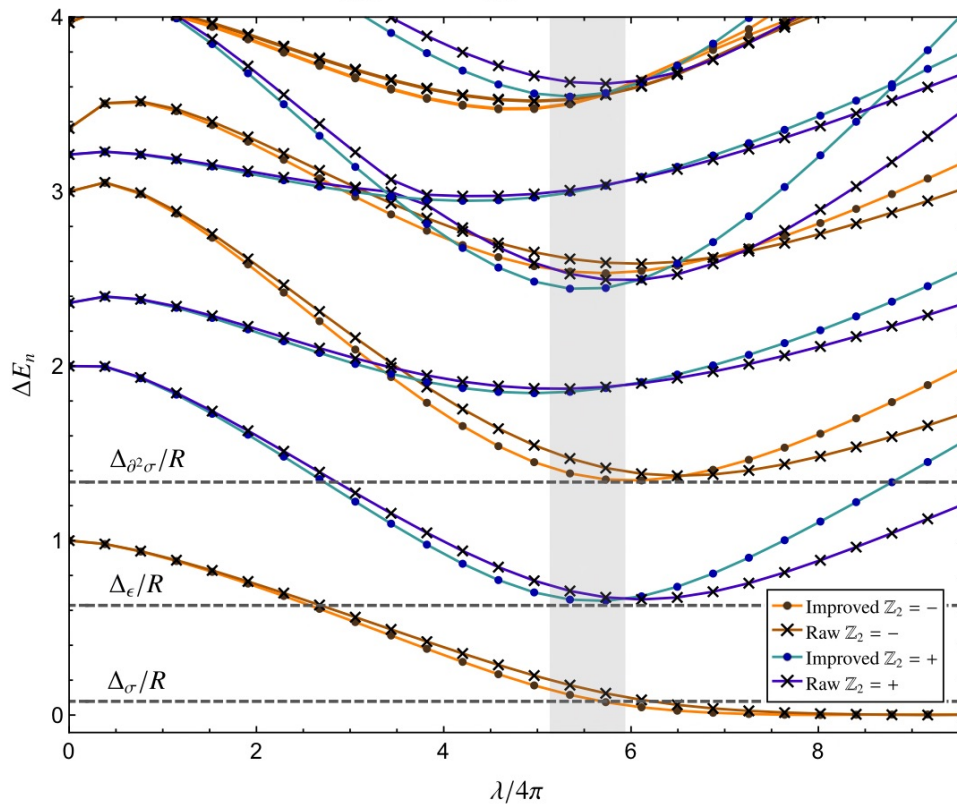


Numerical Results



Numerical Results

$$E_{\max} = 27, m_{\text{NO}} = 1, 2\pi R = 10$$



Stay Tuned

Extension to next order which requires incorporating state dependence into matching coefficients.

Extension to 3D $\lambda\phi^4$ requires incorporating UV divergences

Outlook

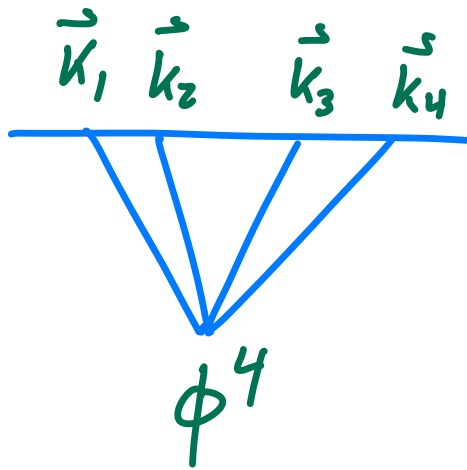
Effective Field Theory
is everywhere...

- Heavy physics decouples
- EFT is dimensional analysis and Taylor expansions
- One loop matching is "solved" using functional methods
- Stochastic Inflation is EFT Renormalization Group evolution
- Hamiltonian Truncation is EFT with a finite energy cutoff

Backups

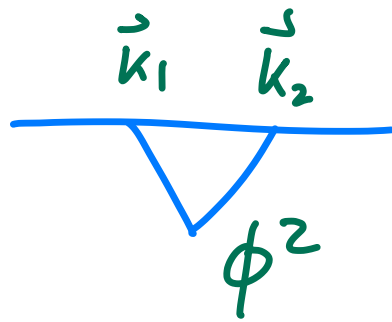
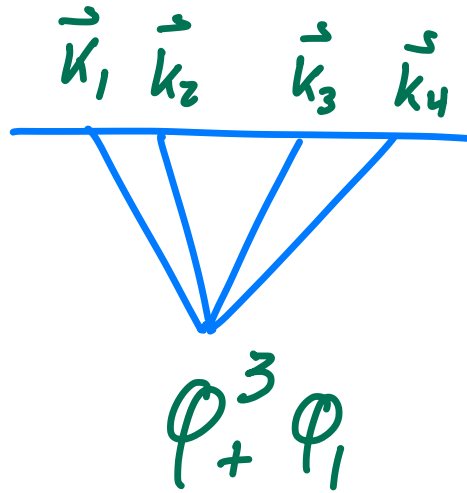
Tree Matching

Assume UV theory is $\lambda \phi^4$



$=$

$\lambda = \mathcal{L}_{3,1}$

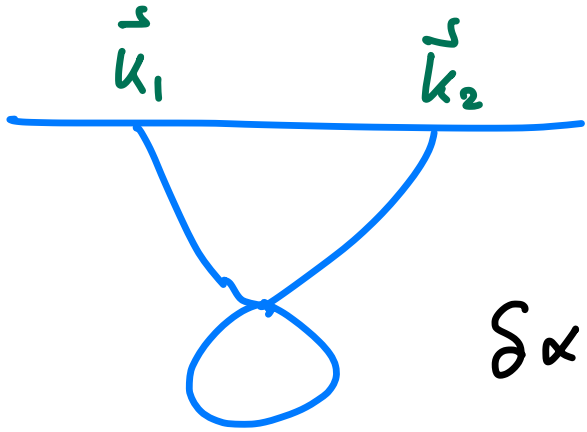


Initial Conditions

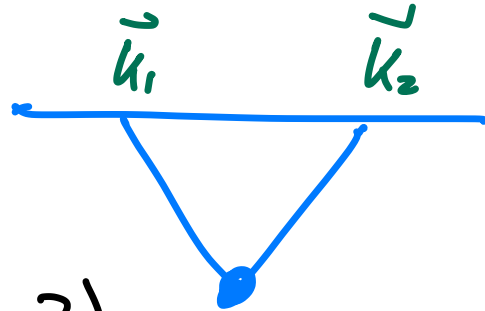
$=$

$\langle \phi_+(k_1) \phi_+(k_2) \rangle_{IC}^{\text{Gauss}}$

One Loop Matching



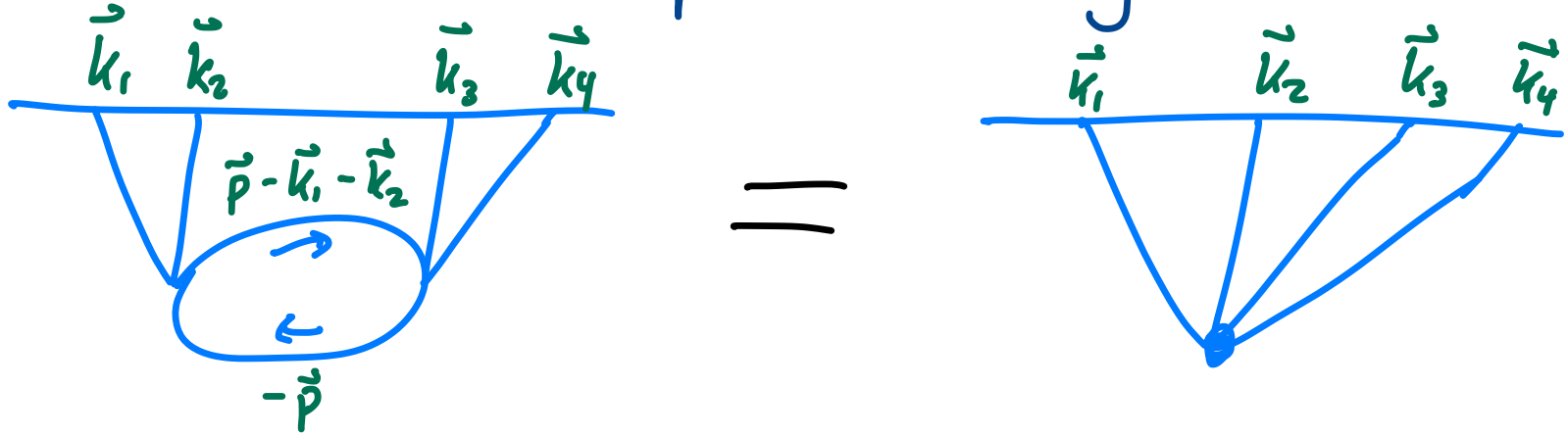
=



$$\delta\alpha = \frac{\lambda}{8\pi^2} \frac{1}{3} \left(\gamma_E - \frac{7}{3} \right)$$

Other terms removed by $\varphi_- \rightarrow \varphi_- + \frac{\lambda}{g} (a_H)^3 \varphi_+^2$

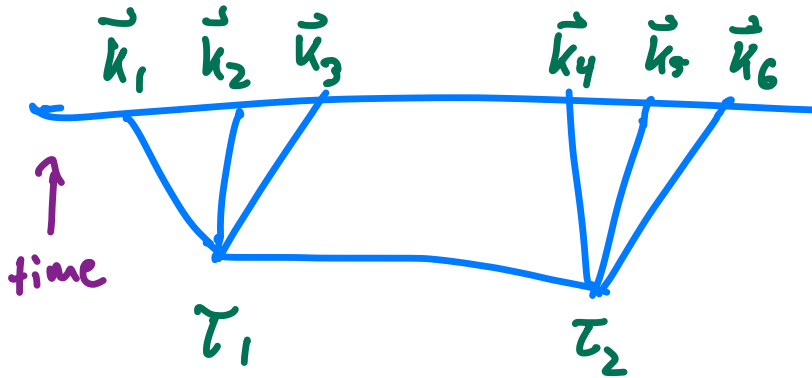
One Loop Matching



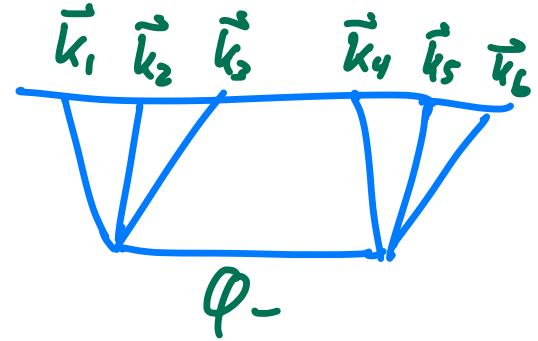
$$C_{3,1} = \lambda - \frac{\lambda^2}{4\pi^2} \left(\frac{1}{9} \gamma_E (2 + 3\gamma_E) + \frac{5}{12} \pi^2 \right)$$

+ $\mathcal{O}(\lambda^2)$ impact on initial conditions
(contributes to $NNLO$ RG)

Initial Conditions to λ^2

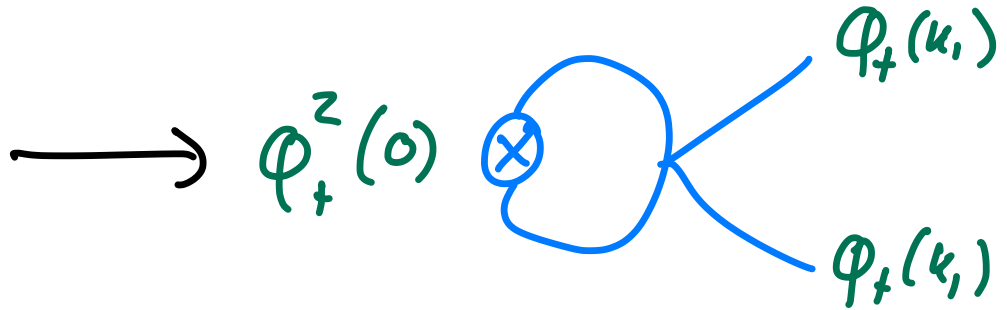
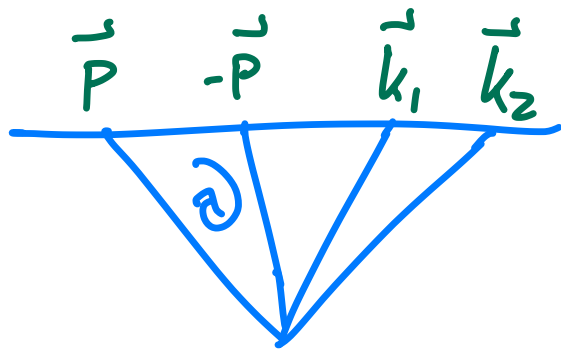


$=$
 $\mathcal{O}(\lambda^2)$



$+ \langle \varphi_+(\vec{k}_1) \dots \varphi_+(\vec{k}_6) \rangle_{IC}$

NLO $\phi_+^2(0)$ RG

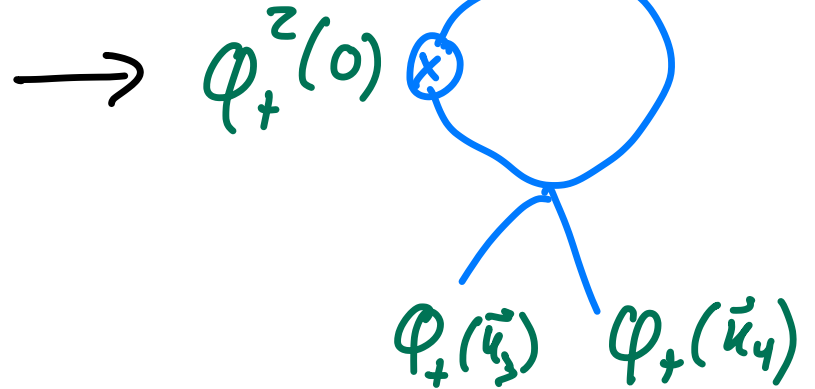
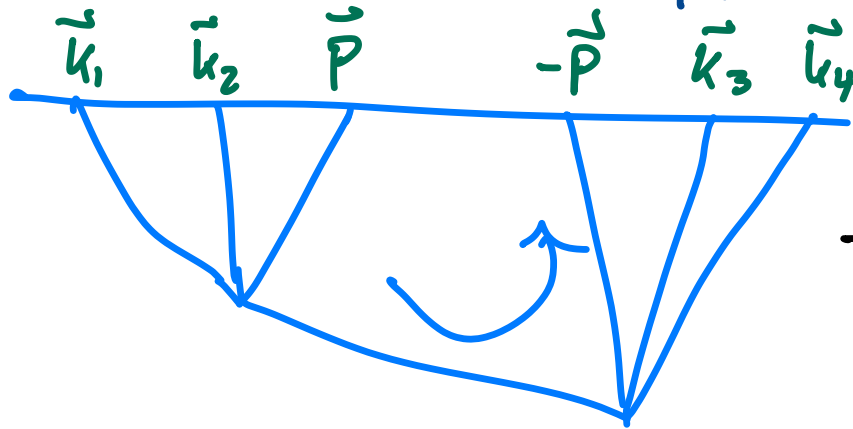


$$\langle \phi_+^2[\vec{x}=0] \phi_+(\vec{k}_1) \phi_+(\vec{k}_2) \rangle$$

$$= \lambda P(k_1) P(k_2) \left(\frac{1}{48\pi^2 \alpha^2} + \frac{(4 - 3\gamma_E - 3\log \bar{\mu})}{72\pi^2 \alpha} + \text{finite} \right)$$

← $\pm R$

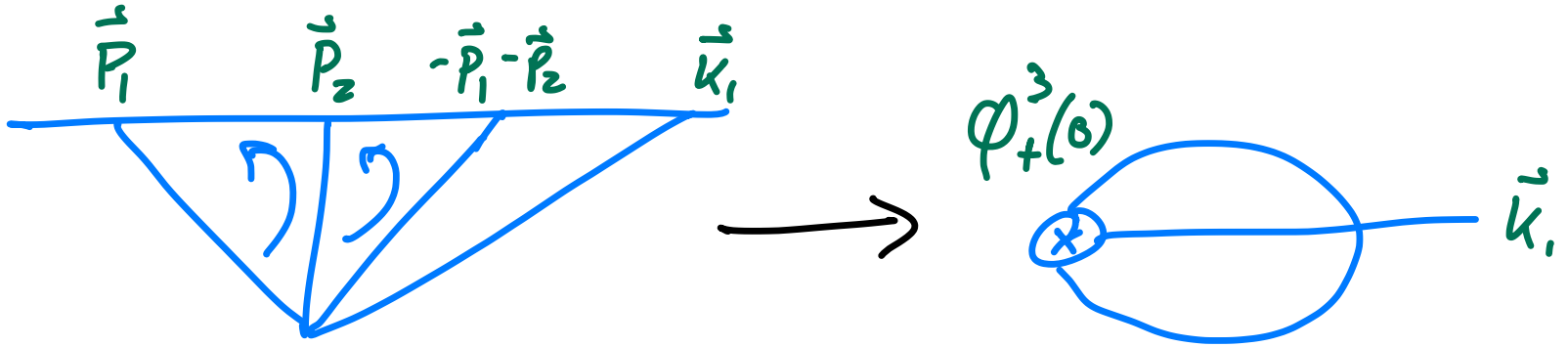
NNLO $\varphi_+^2(0)$ RG



$$\langle \varphi_+^2[\vec{x}=0] \varphi_+(k_1) \dots \varphi_+(k_4) \rangle \sim$$

$$\frac{\lambda^2}{8\pi^2\alpha} \frac{1}{27} \left[16 + 4\gamma_E(-11 + 3\gamma_E) + 3\pi^2 + 12(\log 2)^2 \right] + \dots$$

NNLO $\phi^3(0)$ RG



$$\langle \phi_+^3[\vec{x}=0] \phi_+(\vec{k}_1) \rangle = \frac{\lambda}{16\pi^2} \frac{1}{12} P(k_1) \left[\frac{1}{\alpha} + \dots \right]$$

Put it all together

$$\frac{\partial}{\partial t} P = \frac{1}{3} \frac{\partial}{\partial \phi_+} [V_{\text{eff}}' P] + \frac{1}{8\pi^2} \frac{\partial^2}{\partial \phi_+^2} P + \frac{\lambda_{\text{eff}}}{192\pi^2} \frac{\partial^3}{\partial \phi_+^3} (P_+ P)$$

$$V_{\text{eff}} = \frac{\lambda_{\text{eff}}}{3!} \left(\phi_+^3 + \frac{\lambda_{\text{eff}}}{18} \phi_+^5 + \frac{\lambda_{\text{eff}}}{162} \phi_+^7 + \dots \right)$$

$$\lambda_{\text{eff}} = \lambda - 12b_2 - \frac{\lambda^2}{2\pi^2} \left(\frac{1}{3} \gamma_E (2 + 3\gamma_E) + \frac{5\pi^2}{4} \right)$$

Compute equilibrium distribution,
relaxation eigenvalues, etc.