

Global fits to $b \rightarrow c\tau\nu$ transitions

Mainz Prisma U.

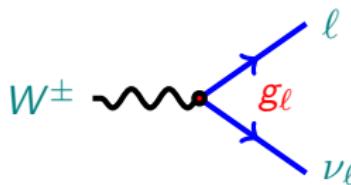
17th November 2020

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A. Pich and M. Jung

Based on arXiv:1904.09311 and arXiv:2004.06726

Introduction

Lepton Flavour Universality \rightarrow no difference between electroweak decays at high energy (negligible mass)



$$g_e = g_\mu = g_\tau$$

anomalies in
Lepton Universality

charged currents $b \rightarrow c$
neutral currents $b \rightarrow s$

$$\left(\frac{\tau}{\mu, e}, \mathcal{R}_{D^{(*)}} \right)$$

$$\left(\frac{\mu}{e}, \mathcal{R}_{K^{(*)}} \right)$$

Introduction

Discrepancies $[(2 - 4)\sigma]$ in semileptonic B decays

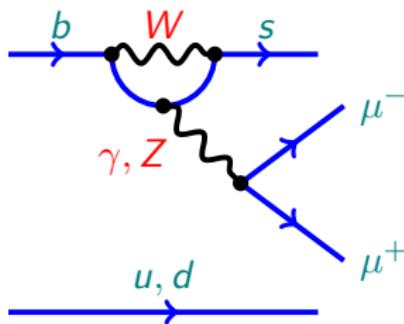
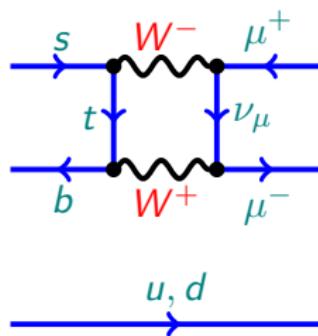
$$\mathcal{R}_{D^{(*)}} \equiv \frac{\mathcal{B}(B \rightarrow D^{(*)}\tau\bar{\nu})}{\mathcal{B}(B \rightarrow D^{(*)}\ell\bar{\nu})} \Big|_{\ell=e,\mu} \quad \mathcal{R}_{D^{(*)}}^{\text{exp}} > \mathcal{R}_{D^{(*)}}^{\text{SM}}$$

$$\mathcal{R}_{K^{(*)}} \equiv \frac{\mathcal{B}(B \rightarrow K^{(*)}\mu\mu)}{\mathcal{B}(B \rightarrow K^{(*)}ee)} \Big|_{q^2 \in [q_{\min}^2, q_{\max}^2]} \quad \mathcal{R}_{K^{(*)}}^{\text{exp}} < \mathcal{R}_{K^{(*)}}^{\text{SM}}$$

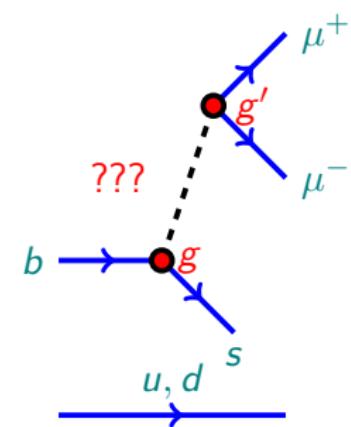
Introduction: $B \rightarrow K\mu^+\mu^-$

$$B = B^-, B^0, \\ K = K^0, K^+,$$

$$B^- = b \bar{u}, B^0 = \bar{b} d \\ K^+ = u \bar{s}, K^0 = d \bar{s}$$

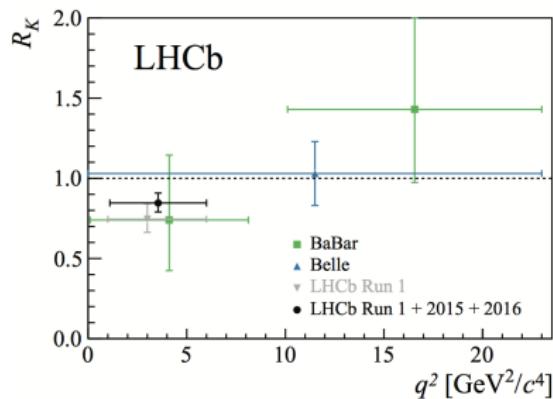
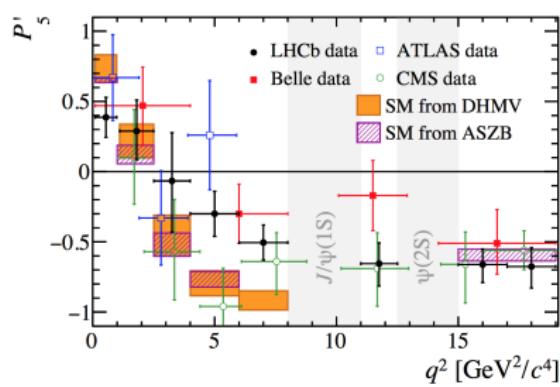


Standard Model



New Physics

Introduction: $B \rightarrow K\mu^+\mu^-$

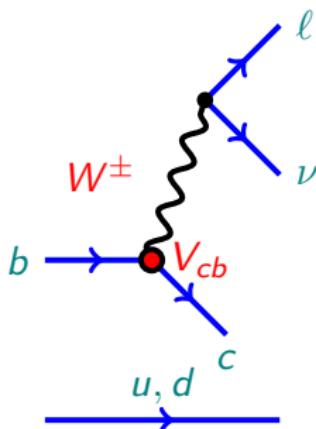


Also $B_d \rightarrow K^0\mu^+\mu^-$, $B_s \rightarrow \phi\mu^+\mu^-$, $\Lambda_d \rightarrow \Lambda\mu^+\mu^- \dots$

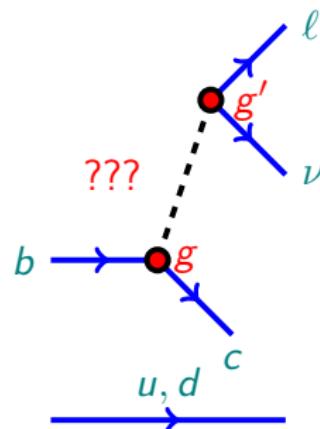
Introduction: $B \rightarrow D\tau\nu$

$$B = B^-, B^0, \\ D = D^0, D^+,$$

$$B^- = b \bar{u}, B^0 = \bar{b}d \\ D^0 = c \bar{u}, D^+ = c \bar{d}$$



Standard Model



New Physics

Current status of charged B anomalies

$$\mathcal{R}_{D^{(*)}} \equiv \frac{\mathcal{B}(B \rightarrow D^{(*)}\tau\bar{\nu})}{\mathcal{B}(B \rightarrow D^{(*)}\ell\bar{\nu})} \quad \begin{array}{ll} 2016 : 3.9\sigma & \text{discrepancy} \\ 2019 : 3.2\sigma & \text{discrepancy} \end{array}$$

SM prediction

$$\mathcal{R}_D^{\text{SM}} = 0.300_{-0.004}^{+0.005} \quad \text{and} \quad \mathcal{R}_{D^*}^{\text{SM}} = 0.251_{-0.003}^{+0.004}$$

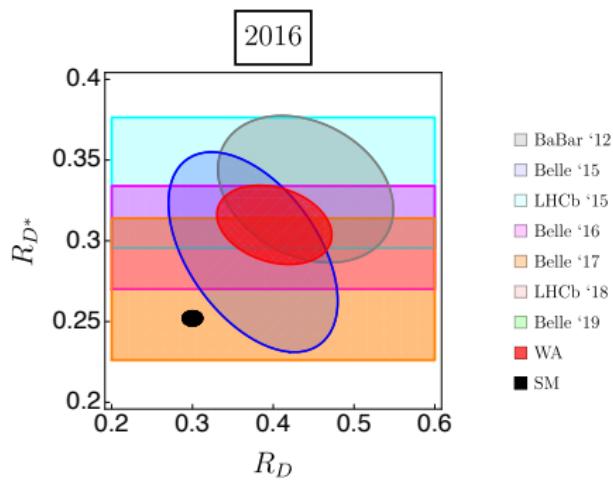
World average 2016 (HFLAV)

$$\mathcal{R}_D = 0.403 \pm 0.040 \pm 0.02 \quad \text{and} \quad \mathcal{R}_{D^*} = 0.310 \pm 0.015 \pm 0.008$$

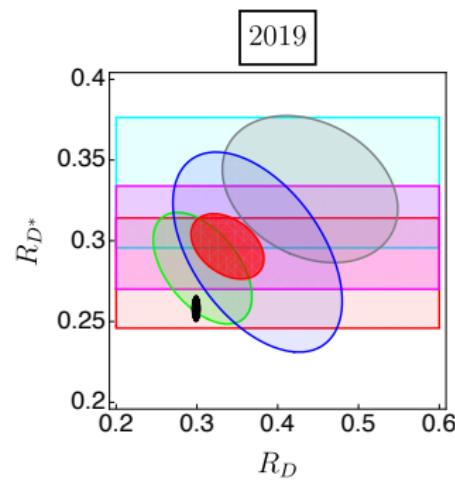
World average 2019 (HFLAV)

$$\mathcal{R}_D = 0.340 \pm 0.027 \pm 0.013 \quad \text{and} \quad \mathcal{R}_{D^*} = 0.295 \pm 0.011 \pm 0.008$$

Current status of charged B anomalies



World average 2016 (HFLAV)



World average 2019 (HFLAV)

Current status of charged B anomalies

More measurements...

$$\bar{F}_L^{D^*} = 0.60 \pm 0.08 \text{ (stat)} \pm 0.04 \text{ (syst)} \quad 1.6\sigma \text{ discrepancy}$$

[Belle '19]

Current status of charged B anomalies

More measurements...

$$\bar{F}_L^{D^*} = 0.60 \pm 0.08 \text{ (stat)} \pm 0.04 \text{ (syst)} \quad 1.6\sigma \text{ discrepancy}$$

[Belle '19]

$$\mathcal{R}_{J/\psi} \equiv \frac{\mathcal{B}(B_c \rightarrow J/\psi \tau \bar{\nu}_\tau)}{\mathcal{B}(B_c \rightarrow J/\psi \mu \bar{\nu}_\mu)} = 0.71 \pm 0.17 \pm 0.18, \quad \mathcal{R}_{J/\psi}^{SM} \approx 0.25 - 0.28$$

before October!

$$\mathcal{P}_\tau^{D^*} = -0.38 \pm 0.51^{+0.21}_{-0.16}, \quad \mathcal{P}_\tau^{D^*SM} = -0.497 \pm 0.013$$

...

Large experimental uncertainties

Current status of charged B anomalies

More measurements...

$$\bar{F}_L^{D^*} = 0.60 \pm 0.08 \text{ (stat)} \pm 0.04 \text{ (syst)} \quad 1.6\sigma \text{ discrepancy}$$

[Belle '19]

$$\mathcal{R}_{J/\psi} \equiv \frac{\mathcal{B}(B_c \rightarrow J/\psi \tau \bar{\nu}_\tau)}{\mathcal{B}(B_c \rightarrow J/\psi \mu \bar{\nu}_\mu)} = 0.71 \pm 0.17 \pm 0.18, \quad \mathcal{R}_{J/\psi}^{SM} = 0.2582 \pm 0.0038$$

[2007.06956]

$$\mathcal{P}_\tau^{D^*} = -0.38 \pm 0.51^{+0.21}_{-0.16}, \quad \mathcal{P}_\tau^{D^*SM} = -0.497 \pm 0.013$$

...

Large experimental uncertainties

Theoretical framework - Effective Hamiltonian

- Most general $SU(3)_C \otimes U(1)_Q$ -invariant effective Hamiltonian at b scale, without light right-handed neutrinos

$$\mathcal{H}_{\text{eff}}^{b \rightarrow c \ell \nu} = \frac{4G_F}{\sqrt{2}} V_{cb} \left[(1 + C_{LL}^V) \mathcal{O}_{LL}^V + C_{RL}^V \mathcal{O}_{RL}^V + C_{RL}^S \mathcal{O}_{RL}^S + C_{LL}^S \mathcal{O}_{LL}^S + C_{LL}^T \mathcal{O}_{LL}^T \right] + \text{h.c.}$$

$$\mathcal{O}_{LL}^V = (\bar{c}_L \gamma^\mu b_L) (\bar{\ell}_L \gamma_\mu \nu_{\ell L}),$$

$$\mathcal{O}_{RL}^V = (\bar{c}_R \gamma^\mu b_R) (\bar{\ell}_L \gamma_\mu \nu_{\ell L}),$$

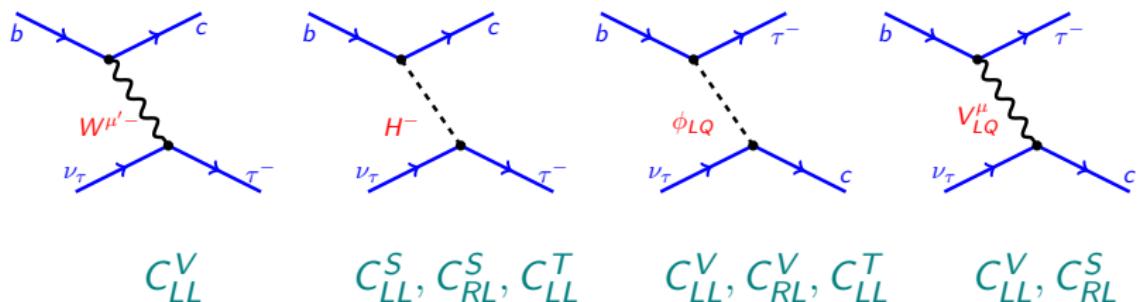
$$\mathcal{O}_{RL}^S = (\bar{c}_L b_R) (\bar{\ell}_R \nu_{\ell L}),$$

$$\mathcal{O}_{LL}^S = (\bar{c}_R b_L) (\bar{\ell}_R \nu_{\ell L}),$$

$$\mathcal{O}_{LL}^T = (\bar{c}_R \sigma^{\mu\nu} b_L) (\bar{\ell}_R \sigma_{\mu\nu} \nu_{\ell L}).$$

$$C_{LL}^{V\text{SM}} = C_{RL}^{V\text{SM}} = C_{LL}^{S\text{SM}} = C_{RL}^{S\text{SM}} = C_{LL}^{T\text{SM}} = 0$$

Theoretical framework - Effective Hamiltonian



Many analysis: usually with single operator/mediator and partial data information

[Freytsis et al, Bardhan et al, Cai et al, Hu et al, Celis et al, Datta et al, Bhattacharya et al, Alonso et al, . . .]

Theoretical framework - Assumptions

- NP contributions, $C_i \neq 0$, only in the third generation of leptons

$\mathcal{B}(\tau \rightarrow \mu \nu_\tau \bar{\nu}_\mu) / \mathcal{B}(\tau \rightarrow e \nu_\tau \bar{\nu}_e) \approx 1$
analyses of $b \rightarrow c(e, \mu) \bar{\nu}_{(e, \mu)}$ transitions

- EWSB is linearly related $\rightarrow C_{RL}^V$ is flavour universal, i.e.
 $C_{RL}^V = 0$

No new states up to 1 TeV
Higgs couplings consistent with SM

- CP-conserving: all Wilson coefficients C_i are assumed to be real

Form Factors

- Heavy quark effective theory (HQET) parametrization
- Corrections of order α_s , $\Lambda_{\text{QCD}}/m_{b,c}$ and $\Lambda_{\text{QCD}}^2/m_c^2$
- Inputs from lattice QCD, light cone sum rules and QCD sum rules
- No experimental information used \rightarrow FFs independent of NP scenario
- 10 form-factor parameters

$$\hat{h}(q^2) = h(q^2)/\xi(q^2).$$

[M. Jung et. al '18, M. Bordrone et al, '19]

Form Factors

Parameter	Value	
ρ^2	1.32 ± 0.06	
c	1.20 ± 0.12	
d	-0.84 ± 0.17	
$\chi_2(1)$	-0.058 ± 0.020	
$\chi'_2(1)$	0.001 ± 0.020	
$\chi'_3(1)$	0.036 ± 0.020	
$\eta(1)$	0.355 ± 0.040	
$\eta'(1)$	-0.03 ± 0.11	
$l_1(1)$	0.14 ± 0.23	
$l_2(1)$	-2.00 ± 0.30	
		$\left. \right\} \text{leading IW function}$
		$\left. \right\} \mathcal{O}(1/m_{b,c})$
		$\left. \right\} \mathcal{O}(1/m_c^2)$

$$\xi(q^2) = 1 - 8\rho^2 z(q^2) + (64c - 16\rho^2) z^2(q^2) + (256c - 24\rho^2 + 512d) z^3(q^2)$$

Observables in the fit

$$\begin{aligned}
 \frac{d\Gamma(\bar{B} \rightarrow D\tau\bar{\nu}_\tau)}{dq^2} &= \frac{G_F^2 |V_{cb}|^2}{192\pi^3 m_B^3} q^2 \sqrt{\lambda_D(q^2)} \left(1 - \frac{m_\tau^2}{q^2}\right)^2 \\
 &\times \left\{ \left| 1 + C_{LL}^V + C_{RL}^V \right|^2 \left[\left(1 + \frac{m_\tau^2}{2q^2}\right) H_{V,0}^{s,2} + \frac{3}{2} \frac{m_\tau^2}{q^2} H_{V,t}^{s2} \right] \right. \\
 &+ \frac{3}{2} \left| C_{RL}^S + C_{LL}^S \right|^2 H_S^s + 8 |C_T|^2 \left(1 + \frac{2m_\tau^2}{q^2}\right) H_T^{s2} \\
 &+ 3 \operatorname{Re} \left[\left(1 + C_{LL}^V + C_{RL}^V\right) \left(C_{RL}^S + C_{LL}^S\right)^* \right] \frac{m_\tau}{\sqrt{q^2}} H_S^s H_{V,t}^s \\
 &- \left. 12 \operatorname{Re} \left[\left(1 + C_{LL}^V + C_{RL}^V\right) \left(C_{LL}^T\right)^* \right] \frac{m_\tau}{\sqrt{q^2}} H_T^s H_{V,0}^s \right\}
 \end{aligned}$$

vector contribution $C_V = 1 + C_{LL}^V + C_{RL}^V$

scalar contribution $C_S = C_{RL}^S + C_{LL}^S$

Observables in the fit

$$\begin{aligned}
 \frac{d\Gamma(\bar{B} \rightarrow D^* \tau \bar{\nu}_\tau)}{dq^2} = & \frac{G_F^2 |V_{cb}|^2}{192\pi^3 m_B^3} q^2 \sqrt{\lambda_{D^*}(q^2)} \left(1 - \frac{m_\tau^2}{q^2}\right)^2 \\
 \times & \left\{ \left(\left| 1 + (C_{LL}^V)^2 + |C_{RL}^V|^2 \right|^2 \right) \left[\left(1 + \frac{m_\tau^2}{2q^2}\right) (H_{V,+}^2 + H_{V,-}^2 + H_{V,0}^2) + \frac{3}{2} \frac{m_\tau^2}{q^2} H_{V,t}^2 \right] \right. \\
 - & 2 \operatorname{Re} \left[(1 + C_{LL}^V) (C_{RL}^V)^* \right] \left[\left(1 + \frac{m_\tau^2}{2q^2}\right) (H_{V,0}^2 + 2H_{V,+}H_{V,-}) + \frac{3}{2} \frac{m_\tau^2}{q^2} H_{V,t}^2 \right] \\
 + & \frac{3}{2} \left| C_{RL}^S - C_{LL}^S \right|^2 H_S^2 + 8 |C_T|^2 \left(1 + \frac{2m_\tau^2}{q^2}\right) (H_{T,+}^2 + H_{T,-}^2 + H_{T,0}^2) \\
 + & 3 \operatorname{Re} \left[(1 + C_{LL}^V - C_{RL}^V) (C_{RL}^S - C_{LL}^S)^* \right] \frac{m_\tau}{\sqrt{q^2}} H_S H_{V,t} \\
 - & 12 \operatorname{Re} \left[(1 + C_{LL}^V) (C_{LL}^T)^* \right] \frac{m_\tau}{\sqrt{q^2}} (H_{T,0} H_{V,0} + H_{T,+} H_{V,+} - H_{T,-} H_{V,-}) \\
 + & \left. 12 \operatorname{Re} \left[C_{RL}^V (C_{LL}^T)^* \right] \frac{m_\tau}{\sqrt{q^2}} (H_{T,0} H_{V,0} + H_{T,+} H_{V,-} - H_{T,-} H_{V,+}) \right\}
 \end{aligned}$$

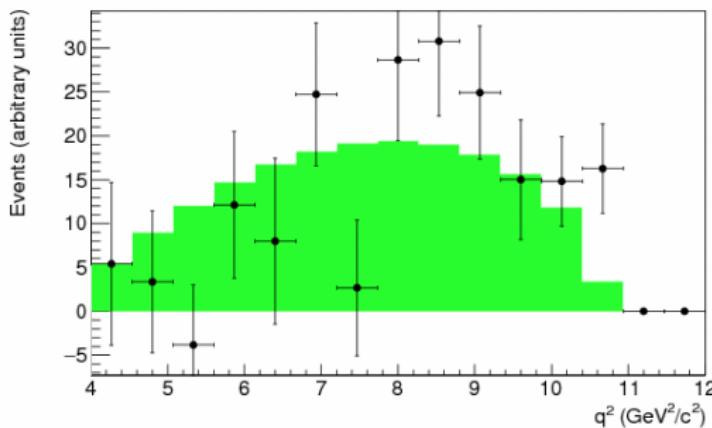
"almost" axial contribution $C_A = C_{RL}^V - (1 + C_{LL}^V)$

pseudoscalar contribution $C_P = C_{RL}^S - C_{LL}^S$

Observables in the fit

Observables included in the fit:

- The ratios $\mathcal{R}_{D^{(*)}}$
- Differential distributions of the decay rates $\Gamma(B \rightarrow D^{(*)}\tau\bar{\nu}_\tau)$
- The longitudinal polarization fraction $\bar{F}_L^{D^*}$



[Belle '15]

Observables in the fit

Observables included in the fit:

- The ratios $\mathcal{R}_{D^{(*)}}$
- Differential distributions of the decay rates $\Gamma(B \rightarrow D^{(*)}\tau\bar{\nu}_\tau)$
- The longitudinal polarization fraction $\bar{F}_L^{D^*}$

$$F_L^{D^*}(q^2) = \frac{d\Gamma_{\lambda_{D^*}=0}}{dq^2} \Bigg/ \frac{d\Gamma^{D^*}}{dq^2}$$

Theoretical framework - Observables in the fit

- The leptonic decay rate $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) \leq 10(30)\%$

$$\begin{aligned} \mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) = \tau_{B_c} & \frac{m_{B_c} m_\tau^2 f_{B_c}^2 G_F^2 |V_{cb}|^2}{8\pi} \left(1 - \frac{m_\tau^2}{m_{B_c}^2}\right)^2 \\ & \times \left| (1 + C_{LL}^V) - C_{RL}^V + \frac{m_{B_c}^2}{m_\tau(m_b + m_c)} (C_{RL}^S - C_{LL}^S) \right|^2 \end{aligned}$$

$\mathcal{B} < 10\%$ LEP data at the Z peak

$\mathcal{B} < 30\%$ B_c lifetime

See also

M. Blanke et. al '18
for further discussion

Fit

$$\chi^2 = \chi_{\text{exp}}^2 + \chi_{\text{FF}}^2$$

$\chi_{\text{exp}}^2 \rightarrow$ experimental contributions: 2 + 58 + 1 observables

$\chi_{\text{FF}}^2 \rightarrow$ 10 form factors

$$\chi^2(y_i) = F(y_i)^T V^{-1} F(y_i), \quad F(y_i) = f_{\text{th}}(y_i) - f_{\text{exp}}, \quad V_{ij} = \rho_{ij} \sigma_i \sigma_j$$

$y_i \rightarrow$ input parameters of the fit

$\rho_{ij} \rightarrow$ correlation between observable i and j

$\sigma_i \rightarrow$ uncertainty of observable *i*

Δy_i : determined minimizing $\chi^2|_{y_i^{\min} + \Delta y_i}$ varying all parameters
that increase $\Delta\chi^2 = 1$

SM fit

- SM fit, $C_i = 0$

$$\chi^2_{\min}/\text{d.o.f.} = 65.5/57 \rightarrow \text{CL of } \sim 20\%$$

- Uncertainties in $d\Gamma/dq^2$ maximally conservative

$$\chi^2_{\min, d\Gamma}/\text{d.o.f.} \sim 43/54$$

- \mathcal{R}_D and \mathcal{R}_{D^*}

$$\chi^2/\text{d.o.f} = 22.6/2 \rightarrow 4.4\sigma\text{-tension}$$

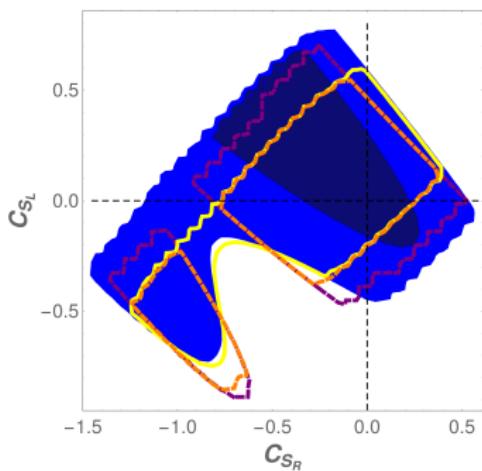
→ NP scenarios judged by the improvement when compared to the SM

Fit and results

- Always two minima with degenerate χ^2

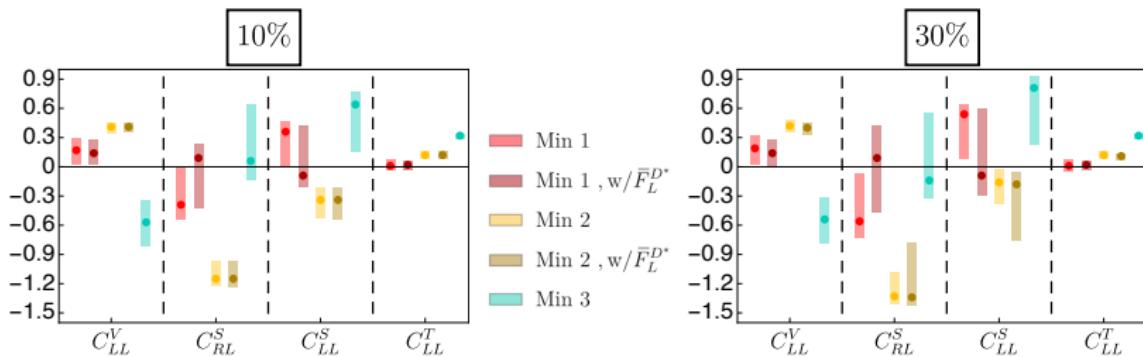
$$C_{LL}^{V'} = -2 - C_{LL}^V \quad \text{and} \quad C_i' = -C_i \quad \text{for } i = S_{RL}, S_{LL}, T_{LL}.$$

- Distributions highly non-gaussian



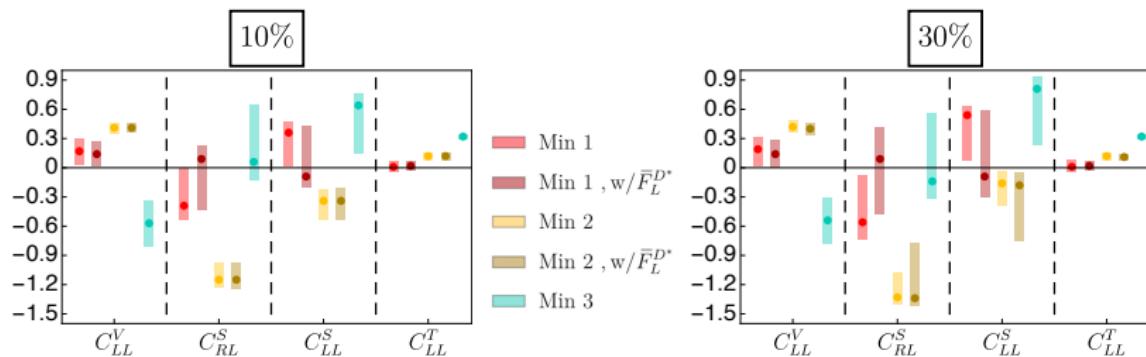
Example of the correlation between two Wilson coefficients ($C_{RL}^S - C_{LL}^S$). Blue areas (lighter 95% and darker 68% CL) minima without $\bar{F}_L^{D^*}$ and $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) \leq 10\%$.

Fit and results



$\mathcal{B}(B_c \rightarrow \tau\nu)$	Min 1	Min 2	Min 3	Min 1	Min 2	Min 3
10%						
$\chi^2_{\min}/\text{d.o.f.}$	34.1/53	37.5/53	58.6/53	33.8/53	36.6/53	58.4/53
C_{LL}^V	$0.17^{+0.13}_{-0.14}$	$0.41^{+0.05}_{-0.06}$	$-0.57^{+0.23}_{-0.24}$	$0.19^{+0.13}_{-0.17}$	$0.42^{+0.06}_{-0.06}$	$-0.54^{+0.23}_{-0.24}$
C_{RL}^S	$-0.39^{+0.38}_{-0.15}$	$-1.15^{+0.18}_{-0.08}$	$0.06^{+0.59}_{-0.19}$	$-0.56^{+0.49}_{-0.17}$	$-1.33^{+0.25}_{-0.08}$	$-0.14^{+0.69}_{-0.18}$
C_{LL}^S	$0.36^{+0.11}_{-0.35}$	$-0.34^{+0.12}_{-0.19}$	$0.64^{+0.13}_{-0.49}$	$0.54^{+0.10}_{-0.46}$	$-0.16^{+0.13}_{-0.22}$	$0.81^{+0.12}_{-0.58}$
C_{LL}^T	$0.01^{+0.06}_{-0.05}$	$0.12^{+0.04}_{-0.04}$	$0.32^{+0.02}_{-0.03}$	$0.01^{+0.07}_{-0.05}$	$0.12^{+0.04}_{-0.04}$	$0.32^{+0.02}_{-0.03}$

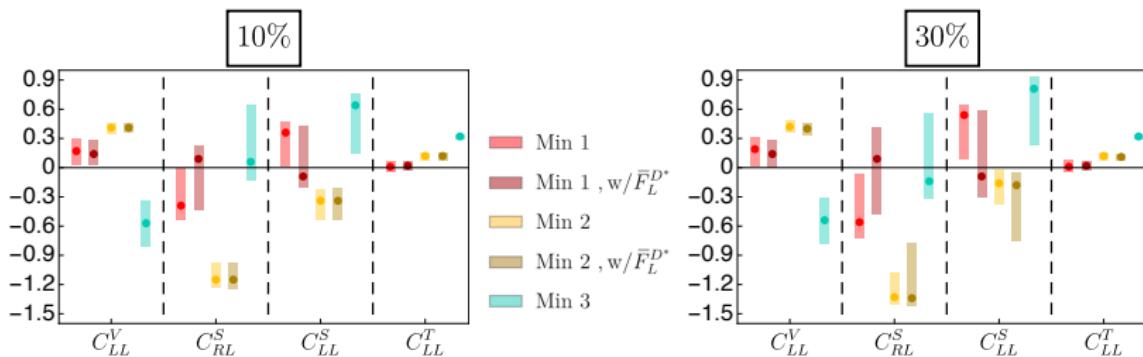
Fit and results



Fits without $\bar{F}_L^{D^*}$

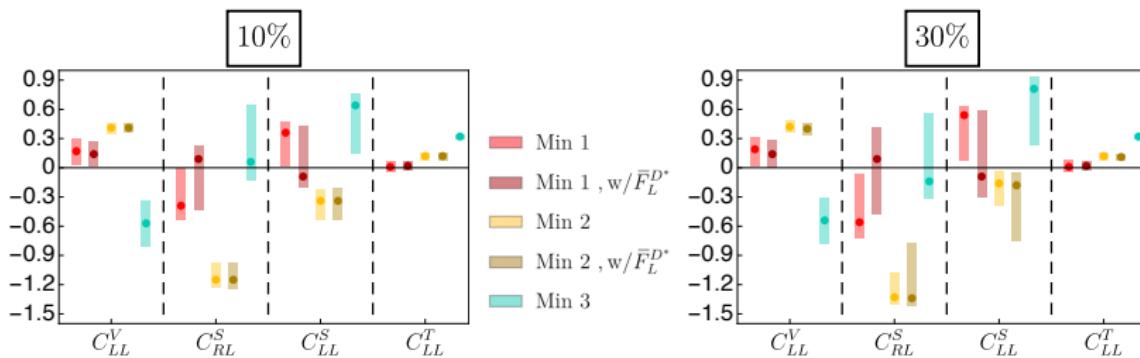
- One global minimum (34.1/53) and two local minima (37.5/53 and 58.6/53)

Fit and results



- Strong preference for New Physics: $\chi_{\text{SM}}^2 - \chi^2 = 31.4$
- All minima saturate the constraint $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) \leq 10 \text{ (30)}\%$
- Complex C_i do not improve the χ^2 , but open to many solutions

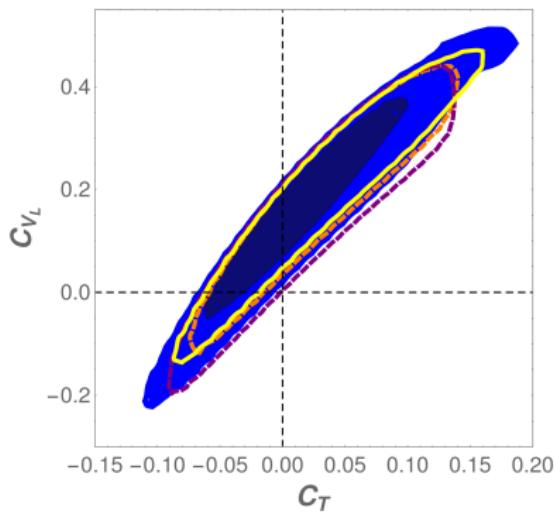
Fit and results



Global minimum (34.1/53)

- No absolute preference of a single Wilson coefficient
- Compatible with a global modification of the SM: adding C_{LL}^V : $\Delta\chi^2 = 1.4$

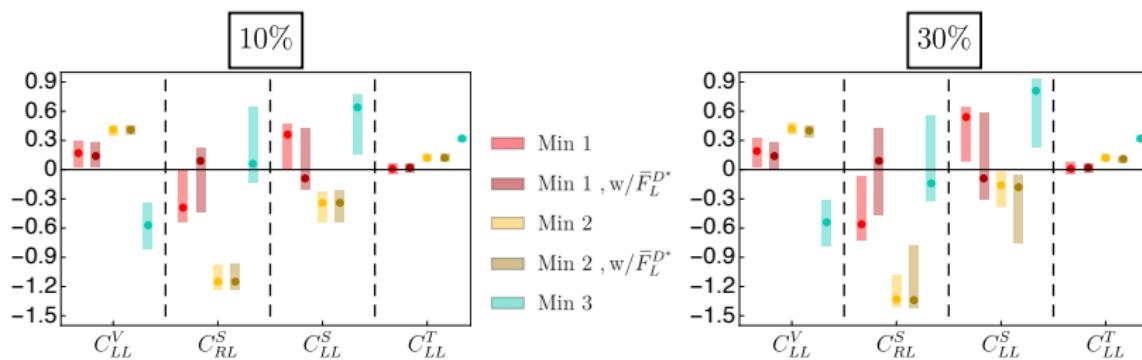
Fit and results



Global minimum (34.1/53)

- Requires either C_{LL}^V or C_{LL}^T

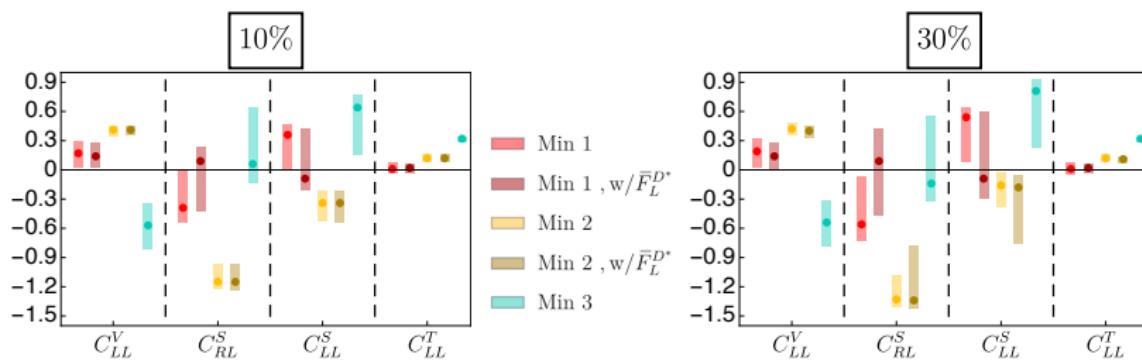
Fit and results



Local minima (Min 2 $\rightarrow 37.5/53$ and Min 3 $\rightarrow 58.6/53$)

- Further away from the SM and involves sizeable Wilson coefficients
- Min 2 fits slightly worse \mathcal{R}_{D^*} and q^2 distributions
- Min 3 fits \mathcal{R}_{D^*} perfectly disfavoured by q^2 distributions

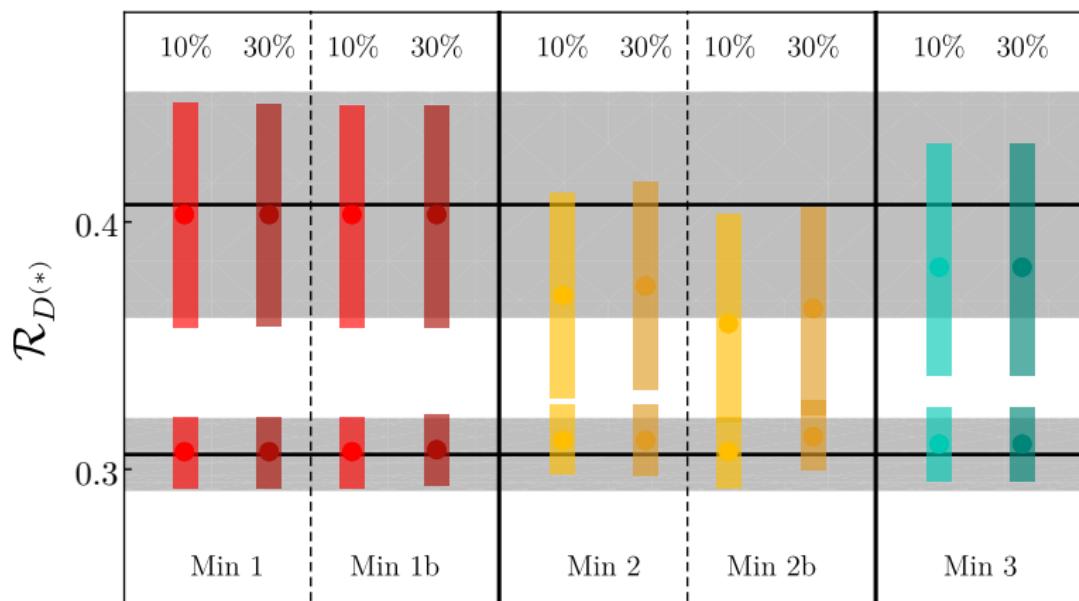
Fit and results



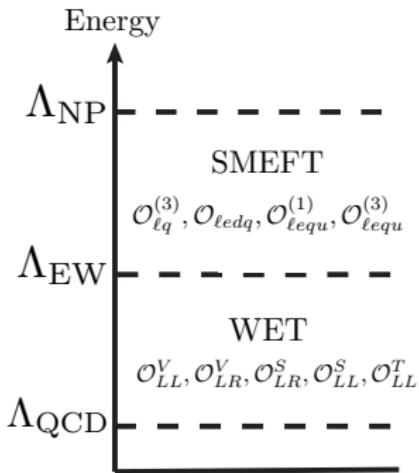
Adding $\bar{F}_L^{D^*}$

- Still no clear preference for a single coefficient
- Central values smaller, while 1σ regions almost constant
- Min 3 disappears

Fit and results



New Physics



$$C_{LL}^V(\mu_b) = -1.503 \tilde{C}_{LL}^V(\Lambda),$$

$$C_{LL}^S(\mu_b) = -1.257 \tilde{C}_{LL}^S(\Lambda) + 0.2076 \tilde{C}_{LL}^T(\Lambda), \quad \mu_b = 5 \text{ GeV}$$

$$C_{RL}^S(\mu_b) = -1.254 \tilde{C}_{RL}^S(\Lambda), \quad \Lambda = 1 \text{ TeV}$$

$$C_{LL}^T(\mu_b) = 0.002725 \tilde{C}_{LL}^S(\Lambda) - 0.6059 \tilde{C}_{LL}^T(\Lambda).$$

New Physics

Spin	Q.N.	Nature	Allowed couplings	SMEFT	WET
0	$S_1 \sim (\bar{3}, 1, 1/3)$	LQ	$\overline{q_L^c} \ell_L, \overline{d_R} u_R^c, \overline{u_R^c} e_R$	$\tilde{C}_{LL}^V, \tilde{C}_{LL}^S, \tilde{C}_{LL}^T$	$C_{LL}^V, C_{LL}^S, C_{LL}^T$
0	$S_3 \sim (\bar{3}, 3, 1/3)$	LQ	$\overline{q_L^c} \ell_L$	\tilde{C}_{LL}^V	C_{LL}^V
0	$R_2 \sim (3, 2, 7/6)$	LQ	$\overline{u_R} \ell_L, \overline{q_L} e_R$	$\tilde{C}_{LL}^S, \tilde{C}_{LL}^T$	C_{LL}^S, C_{LL}^T
0	$H_2 \sim (1, 2, 1/2)$	SB	$\overline{q_L} d_R, \overline{\ell_L} e_R, \overline{u_R} q_L$	$\tilde{C}_{RL}^S, \tilde{C}_{LL}^S$	$C_{RL}^S, C_{LL}^S, C_{LL}^T$
1	$V_2 \sim (\bar{3}, 2, 5/6)$	LQ	$\overline{d_R^c} \gamma_\mu \ell_L, \overline{e_R^c} \gamma_\mu q_L$	\tilde{C}_{RL}^S	C_{RL}^S
1	$U_1 \sim (3, 1, 2/3)$	LQ	$\overline{q_L} \gamma_\mu \ell_L, \overline{d_R} \gamma_\mu e_R$	$\tilde{C}_{LL}^V, \tilde{C}_{RL}^S$	C_{LL}^V, C_{RL}^S
1	$U_3 \sim (3, 3, 2/3)$	LQ	$\overline{q_L} \gamma_\mu \ell_L$	\tilde{C}_{LL}^V	C_{LL}^V
1	$W'_\mu \sim (1, 3, 0)$	VB	$\overline{\ell_L} \gamma_\mu \ell_L, \overline{q_L} \gamma_\mu q_L$	\tilde{C}_{LL}^V	C_{LL}^V

New Physics

Global minimum

$$\begin{aligned} C_{LL}^V &= 0.17^{+0.13}_{-0.14}, & C_{RL}^S &= -0.39^{+0.38}_{-0.15}, \\ C_{LL}^S &= 0.36^{+0.11}_{-0.35}, & C_{LL}^T &= 0.01^{+0.06}_{-0.05}. \end{aligned}$$

$$W'_\mu \sim (1, 3, 0)$$

$$\mathcal{L}_{\text{eff}} \supset -\frac{\tilde{g}_{\ell\nu_\ell}\tilde{g}_{du}^\dagger}{M_{W'}^2} (\bar{\ell}_L \gamma_\mu \nu_{\ell L})(\bar{u}_L \gamma^\mu d_L),$$

$$\frac{M_{W'}}{\left(\tilde{g}_{\ell\nu_\ell}\tilde{g}_{du}^\dagger\right)^{1/2}} \sim 2 \text{ TeV} \xrightarrow[\text{with SM couplings}]{\text{sequential } W'} M_{W'} \sim 0.2 \text{ TeV} \quad \text{ruled out by DS}$$

$$U_3 \sim (3, 3, 1/3), S_3 \sim (\bar{3}, 3, 1/3)$$

New Physics

Min 2

$$C_{LL}^V = 0.41^{+0.05}_{-0.06}, \quad C_{LL}^S = -0.34^{+0.12}_{-0.19},$$

$$C_{RL}^S = -1.15^{+0.18}_{-0.08}, \quad C_{LL}^T = 0.12^{+0.04}_{-0.04}.$$

$$S_1 \sim (\bar{3}, 1, 1/3) + H_2 \sim (1, 2, 1/3)$$

Min 3

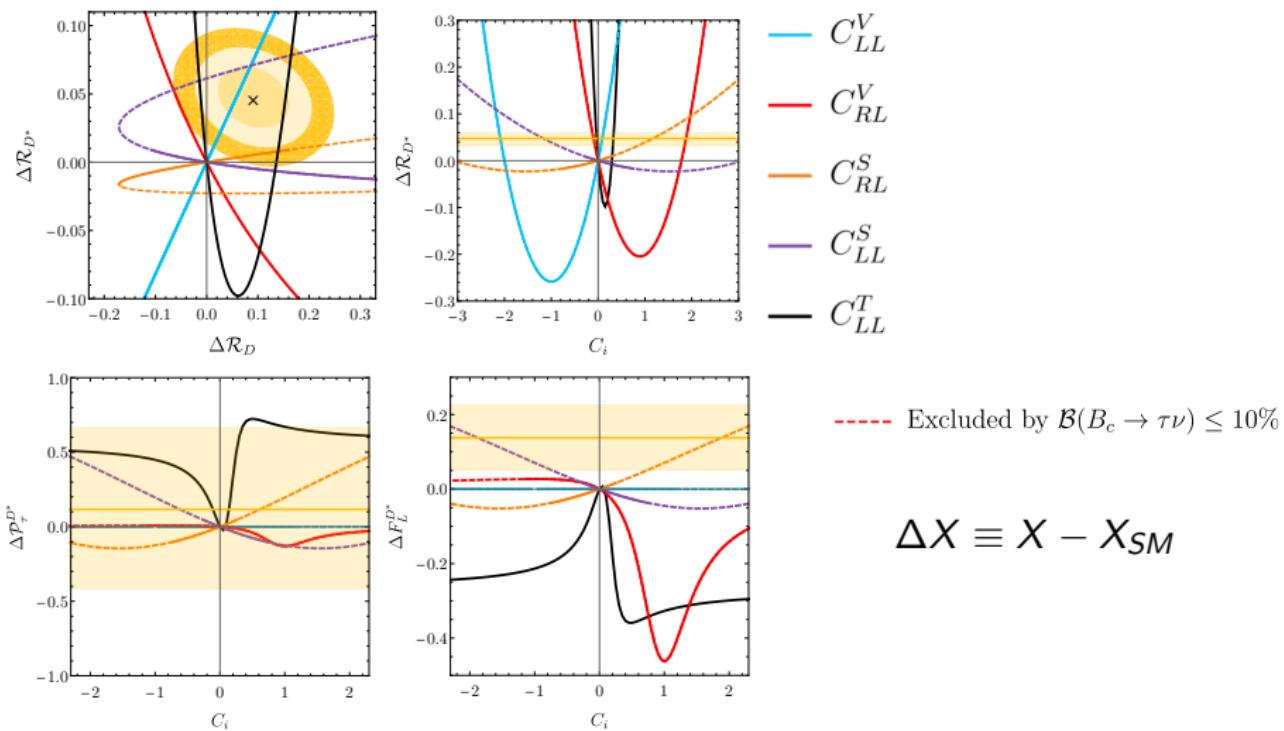
$$C_{LL}^V = -0.57^{+0.23}_{-0.24}, \quad C_{LL}^S = 0.64^{+0.13}_{-0.49},$$

$$C_{RL}^S = 0.06^{+0.59}_{-0.19}, \quad C_{LL}^T = 0.32^{+0.02}_{-0.03}.$$

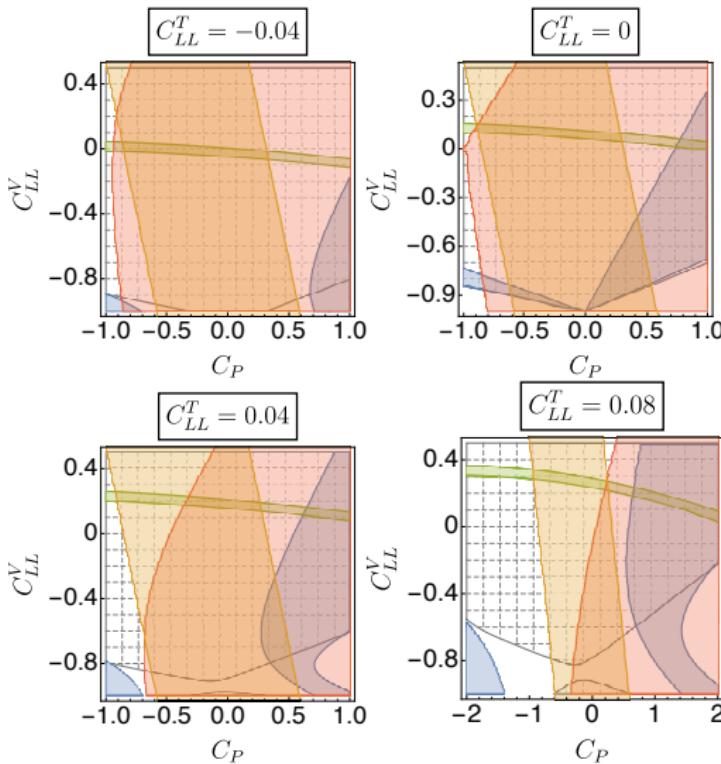
$$R_2 \sim (3, 2, 7/6), S_1 \sim (\bar{3}, 1, 1/3)$$

[A. Angelescu. al '18]

Interpretation of results



Interpretation of results



$$\mathbf{C}_P \equiv \mathbf{C}_{RL}^S - \mathbf{C}_{LL}^S$$

D^* observables =
 $f(C_{LL}^V, C_P, C_T)$

- $\mathcal{P}_\tau^{D^*}$
- $F_L^{D^*}$
- \mathcal{R}_{D^*}
- $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) < 10\%$
- q^2 distribution

It is not possible to accommodate all D^* data at 1σ

Interpretation of results

Not possible to accommodate all experimental data at 1σ

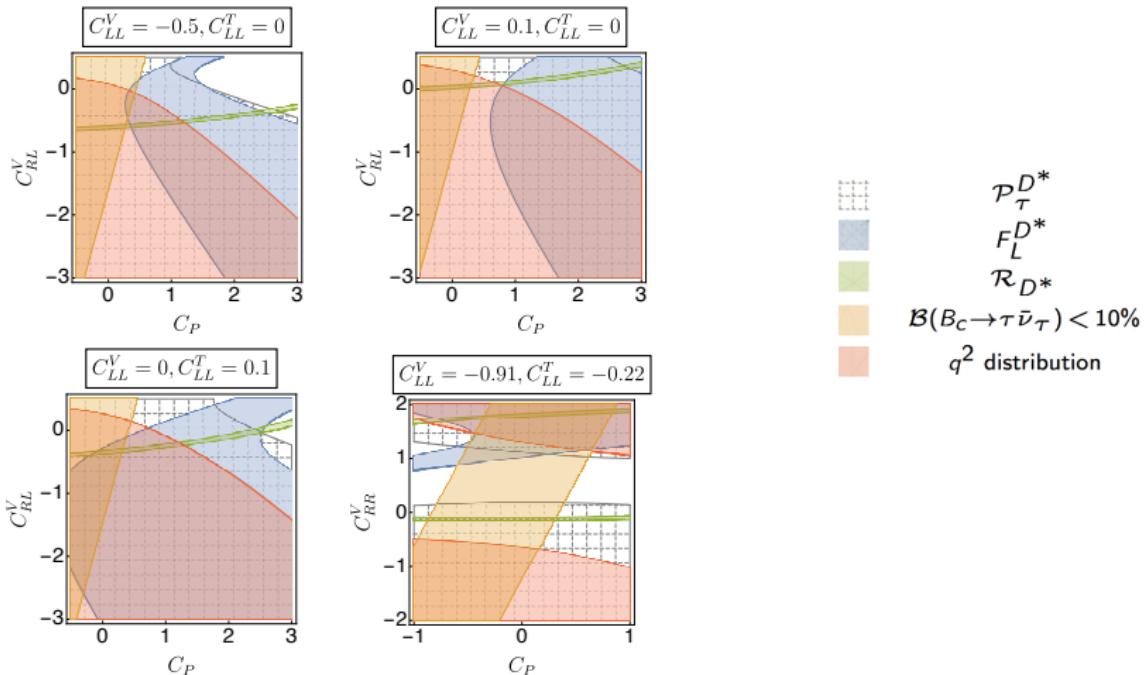
- **Theory side:** one of our assumptions is incorrect
 - There is an insufficient gap between the electroweak and the NP scale
 - The electroweak symmetry breaking is non-linear: C_{RL}^V
 - Additional degrees of freedom: light $\nu_R\dots$
- **Experimental side:** there is an unidentified or underestimated systematic uncertainty in experimental measurements

Interpretation of results

Not possible to accommodate all experimental data at 1σ

- **Theory side:** one of our assumptions is incorrect
 - There is an insufficient gap between the electroweak and the NP scale
 - The electroweak symmetry breaking is non-linear: C_{RL}^V
 - Additional degrees of freedom: light ν_R ...
- **Experimental side:** there is an unidentified or underestimated systematic uncertainty in experimental measurements → upcoming experimental studies of LHCb and Belle II

Adding C_{RL}^V



Adding C_{RL}^V

	Min 4	Min 5	Min 6	Min 7
$\chi^2_{\text{min}}/\text{d.o.f.}$	32.5/53	33.3/53	37.6/53	38.9/53
C_{LL}^V	$-0.91^{+0.10}_{-0.09}$	$-0.85^{+0.20}_{-0.10}$	$0.14^{+0.14}_{-0.12}$	$0.35^{+0.08}_{-0.08}$
C_{RL}^V	$1.89^{+0.19}_{-0.22}$	$-1.58^{+0.23}_{-0.22}$	$0.02^{+0.21}_{-0.24}$	$0.34^{+0.18}_{-0.18}$
C_{RL}^S	$-0.44^{+0.12}_{-0.45}$	$-0.33^{+0.52}_{-0.16}$	$0.10^{+0.15}_{-0.59}$	$-0.68^{+0.54}_{-0.14}$
C_{LL}^S	$-1.34^{+0.49}_{-0.12}$	$0.56^{+0.23}_{-0.54}$	$-0.12^{+0.65}_{-0.15}$	$-0.92^{+0.58}_{-0.11}$
C_{LL}^T	$-0.22^{+0.10}_{-0.11}$	$0.19^{+0.10}_{-0.10}$	$0.01^{+0.09}_{-0.07}$	$-0.02^{+0.08}_{-0.07}$

- Including C_{RL}^V slightly improves the fit $\chi^2/\text{d.o.f.} = 32.5/55$
- Global minimum \rightarrow Min 6
- Two fine-tuned solutions (Min 4 and Min 5): $1 + C_{LL}^V \approx 0$

Including ν_R

$$\mathcal{H}_{\text{eff}} = \frac{4G_F V_{cb}}{\sqrt{2}} \left(\mathcal{O}_{LL}^V + \sum_{\substack{X=S,V,T \\ A,B=L,R}} C_{AB}^X \mathcal{O}_{AB}^X \right),$$

5 Wilson operators \rightarrow 10 Wilson operators

\rightarrow difficult to fit all at the same time
 \rightarrow focus on scenarios

$\bar{P}_\tau^{D^*}$ included in the fits

Spin	Q.N.	Nature	ν_L -WET	ν_R -WET
0	$S_1 \sim (\bar{3}, 1, 1/3)$	LQ	$C_{LL}^V, C_{LL}^S, C_{LL}^T$	$C_{RR}^V, C_{RR}^S, C_{RR}^T$
0	$\Phi \sim (1, 2, 1/2)$	SB	C_{LL}^S, C_{RL}^S	C_{LR}^S, C_{RR}^S
0	$\tilde{R}_2 \sim (3, 2, 1/6)$	LQ	-	C_{RR}^S, C_{RR}^T
1	$U_1^\mu \sim (3, 1, 2/3)$	LQ	C_{LL}^V, C_{RL}^S	C_{RR}^V, C_{LR}^S
1	$\tilde{V}_2^\mu \sim (\bar{3}, 2, -1/6)$	LQ	-	C_{LR}^S
1	$V^\mu \sim (1, 1, -1)$	VB	-	C_{RR}^V

Scenarios

1) RHN + SM-like contribution:

$$\mathcal{O}_{LL}^V, \mathcal{O}_{LR}^V, \mathcal{O}_{RR}^V, \mathcal{O}_{LR}^S, \mathcal{O}_{RR}^S, \mathcal{O}_{RR}^T,$$

2) RHN: $\mathcal{O}_{LR}^V, \mathcal{O}_{RR}^V, \mathcal{O}_{LR}^S, \mathcal{O}_{RR}^S, \mathcal{O}_{RR}^T,$

3) V^μ : $\mathcal{O}_{RR}^V,$

4a) Φ : $\mathcal{O}_{LR}^S, \mathcal{O}_{RR}^S$ (b) + $\mathcal{O}_{LL}^S, \mathcal{O}_{RL}^S$

5a) U_1^μ : $\mathcal{O}_{RR}^V, \mathcal{O}_{LR}^S$, (b) + $\mathcal{O}_{RL}^V, \mathcal{O}_{LL}^S$

6) \tilde{R}_2 : $\mathcal{O}_{RR}^S, \mathcal{O}_{RR}^T$ with $C_{RR}^S = 4r C_{RR}^T,$

7a) S_1 : $\mathcal{O}_{RR}^V, \mathcal{O}_{RR}^S, \mathcal{O}_{RR}^T$ (b) + $\mathcal{O}_{LL}^V, \mathcal{O}_{LL}^S, \mathcal{O}_{LL}^T$ with
 $C_{RR}^S = -4r C_{RR}^T$, (b) and $C_{LL}^S = -4r C_{LL}^T$

8) \tilde{V}_2^μ : $\mathcal{O}_{LR}^S.$

Quality of the fits

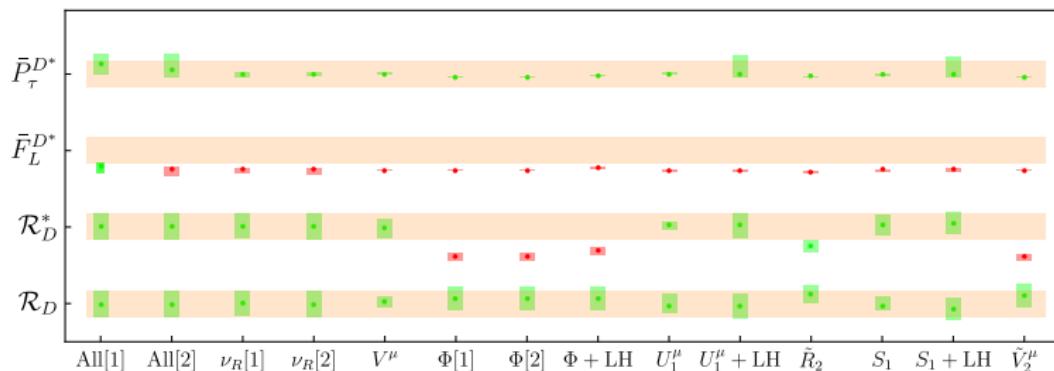
To measure the quality of our fits:

- Comparison with the SM: $\chi^2_{\text{SM}}/\text{d.o.f.} = 52.87/59$
- p -value: Larger p -values correspond to better explanations of the experimental data than lower ones

$$p(\chi^2_{\min}, n) \equiv \int_{\chi^2_{\min}}^{\infty} dz \frac{1}{\chi^2(z, n)} e^{-\chi^2(z, n)}$$

- Pull_{SM}: Compares any fitted solution with the SM results.
Larger pulls \rightarrow further from the SM
- Comparison with experimental data

Fit results



Scenario	$\mathcal{B}(B_c \rightarrow \tau \bar{\nu})$	$\chi^2/\text{d.o.f}$	$\bar{P}_\tau^{D^*}, \bar{F}_L^{D^*}$	Pull_{SM} \mathcal{R}_{D,D^*}	$d\Gamma/dq^2$	Pull_{SM}	$p\text{-value}$
SM	2.16%	52.87/59					69.95%
Scenario 1, Min 1	< 10%	37.26/53	0.007	2.08	0.0414	2.4	95.02%
Scenario 1, Min 2	< 10%	38.86/53	0.001	2.08	0.0006	2.2	92.68%
Scenario 1, Min 1	< 30%	36.42/53	0.022	2.08	0.0866	2.5	96.00%
Scenario 1, Min 2	< 30%	38.54/53	0.011	2.08	0.000	2.2	93.21%
Scenario 2, Min 1	< 10%	38.54/54	0.006	2.32	0.0113	2.5	93.20%
Scenario 2, Min 2	< 10%	39.05/54	0.004	2.32	0.0003	2.4	93.73%
Scenario 2, Min 1	< 30%	38.33/54	0.035	2.32	0.0023	2.5	94.73%
...

Fit results (RHN)

- Difficulty to reproduce longitudinal D^* polarization
- Scenarios involving only scalar (and tensor) operators can't reproduce $\mathcal{R}_{D^*}^{\text{exp}}$ (largest χ^2 value)
- Some scenarios disfavoured by the q^2 differential distributions of the $B \rightarrow D^{(*)}$ decay with respect to the SM (Pull_{SM})
- Clear preference for NP (pulls 1.2 – 3.7)
- V^μ mediator (only C_{RR}^V) → similar to fit without RHN (C_{LL}^V) (largest pull)

Fit results (RHN)

Scenario	$B(B_c \rightarrow \tau \bar{\nu})$	\mathcal{R}_D	\mathcal{R}_{D^*}	$\bar{F}_L^{D^*}$	$\bar{P}_\tau^{D^*}$
Experiment	-	$0.340 \pm 0.027 \pm 0.013$	$0.295 \pm 0.011 \pm 0.008$	$0.60 \pm 0.08 \pm 0.04$	$-0.38 \pm 0.51^{+0.21}_{-0.16}$
<i>Scenario 1, Min 1</i>	10%	0.339 ± 0.030 ✓	0.295 ± 0.014 ✓	$0.494^{+0.025}_{-0.045}$ ✓	$0.06^{+0.43}_{-0.45}$ ✓
<i>Scenario 1, Min 2</i>	10%	0.338 ± 0.030 ✓	0.296 ± 0.014 ✓	$0.472^{+0.023}_{-0.044}$ ✗	$-0.20^{+0.67}_{-0.30}$ ✓
<i>Scenario 1, Min 1</i>	30%	0.338 ± 0.030 ✓	0.295 ± 0.014 ✓	$0.510^{+0.014}_{-0.043}$ ✓	$0.08^{+0.32}_{-0.46}$ ✓
<i>Scenario 1, Min 2</i>	30%	0.338 ± 0.030 ✓	0.296 ± 0.014 ✓	$0.488^{+0.032}_{-0.050}$ ✓	$-0.24^{+0.64}_{-0.28}$ ✓
<i>Scenario 2, Min 1</i>	10%	$0.341^{+0.029}_{-0.028}$ ✓	0.296 ± 0.013 ✓	$0.474^{+0.010}_{-0.024}$ ✗	$-0.42^{+0.13}_{-0.07}$ ✓
<i>Scenario 2, Min 2</i>	10%	0.339 ± 0.030 ✓	0.296 ± 0.014 ✓	$0.471^{+0.012}_{-0.033}$ ✗	$-0.401^{+0.094}_{-0.064}$ ✓
<i>Scenario 2, Min 1</i>	30%	$0.341^{+0.029}_{-0.028}$ ✓	0.296 ± 0.013 ✓	$0.489^{+0.011}_{-0.048}$ ✗	$-0.47^{+0.15}_{-0.05}$ ✓
<i>Scenario 2, Min 2</i>	30%	0.340 ± 0.030 ✓	0.295 ± 0.014 ✓	$0.484^{+0.015}_{-0.045}$ ✗	$-0.45^{+0.13}_{-0.07}$ ✓
<i>Scenario 3</i>	2.5%	0.343 ± 0.012 ✓	0.294 ± 0.010 ✓	0.462 ± 0.004 ✗	$-0.377^{+0.031}_{-0.033}$ ✓
<i>Scenario 4a, Min 1</i>	10%	$0.353^{+0.028}_{-0.027}$ ✓	$0.2638^{+0.0034}_{-0.0049}$ ✗	$0.4662^{+0.0039}_{-0.0057}$ ✗	$-0.5028^{+0.0051}_{-0.0035}$ ✓
<i>Scenario 4a, Min 2</i>	10%	$0.353^{+0.028}_{-0.027}$ ✓	$0.2638^{+0.0034}_{-0.0049}$ ✗	$0.4662^{+0.0039}_{-0.0057}$ ✗	$-0.5028^{+0.0051}_{-0.0034}$ ✓
<i>Scenario 4a, Min 1</i>	30%	$0.348^{+0.028}_{-0.027}$ ✓	$0.2699^{+0.0032}_{-0.0058}$ ✗	$0.4792^{+0.0041}_{-0.0064}$ ✗	$-0.5144^{+0.0056}_{-0.0032}$ ✓
<i>Scenario 4a, Min 2</i>	30%	$0.348^{+0.028}_{-0.027}$ ✓	$0.2699^{+0.0032}_{-0.0058}$ ✗	$0.4792^{+0.0041}_{-0.0064}$ ✗	$-0.5144^{+0.0056}_{-0.0032}$ ✓
<i>Scenario 4b</i>	10%	0.353 ± 0.028 ✓	$0.2708^{+0.0032}_{-0.0052}$ ✗	$0.4815^{+0.0041}_{-0.0068}$ ✗	$-0.442^{+0.005}_{-0.026}$ ✓
<i>Scenario 4b</i>	30%	0.340 ± 0.028 ✓	$0.2866^{+0.0030}_{-0.0081}$ ✓	$0.5125^{+0.0044}_{-0.0126}$ ✓	$-0.356^{+0.006}_{-0.066}$ ✓
<i>Scenario 5a</i>	2.2%	$0.335^{+0.027}_{-0.017}$ ✓	$0.2966^{+0.0043}_{-0.0042}$ ✓	$0.4611^{+0.0056}_{-0.0070}$ ✗	$-0.364^{+0.048}_{-0.050}$ ✓
<i>Scenario 5b</i>	2.0%	0.334 ± 0.029 ✓	0.297 ± 0.013 ✓	$0.4609^{+0.0059}_{-0.0083}$ ✗	$-0.38^{+0.77}_{-0.16}$ ✓
<i>Scenario 6</i>	7.6%	$0.361^{+0.022}_{-0.021}$ ✓	$0.2748^{+0.0066}_{-0.0059}$ ✓	0.4522 ± 0.0050 ✗	$-0.4800^{+0.0078}_{-0.0076}$ ✓
<i>Scenario 7a</i>	4.6%	$0.335^{+0.021}_{-0.011}$ ✓	0.297 ± 0.011 ✓	$0.468^{+0.007}_{-0.011}$ ✗	$-0.377^{+0.033}_{-0.058}$ ✓
<i>Scenario 7b</i>	4.3%	$0.328^{+0.026}_{-0.025}$ ✓	0.299 ± 0.012 ✓	$0.471^{+0.014}_{-0.013}$ ✗	$-0.38^{+0.77}_{-0.12}$ ✓
<i>Scenario 8</i>	7.3%	$0.359^{+0.028}_{-0.027}$ ✓	0.2629 ± 0.0036 ✗	0.4644 ± 0.0043 ✗	-0.5012 ± 0.0039 ✓

Fit results (RHN)

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- Clear preference for NP (pulls $1.2 - 3.7$)
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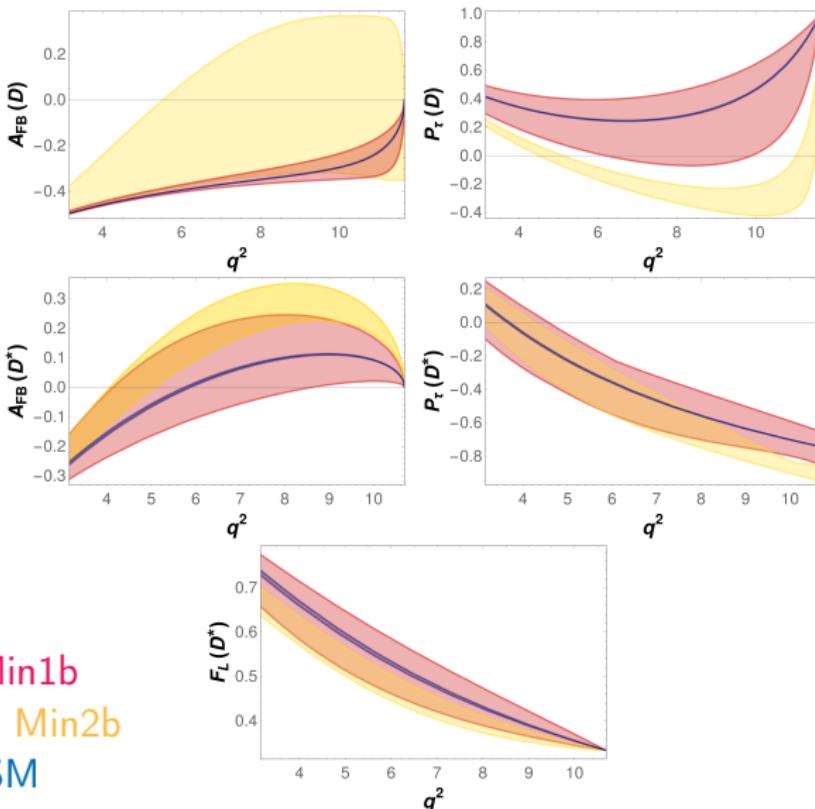
Fit results (RHN)

Scenario	$\mathcal{B}(B_c \rightarrow \tau \bar{\nu})$	$\chi^2/\text{d.o.f}$	$\bar{\mathcal{P}}_{\tau}^{D^*}, F_L^{D^*}$	Pull _{SM} \mathcal{R}_{D,D^*}	$d\Gamma/dq^2$	Pull _{SM}	p-value
SM	2.16%	52.87/59					69.95%
<i>Scenario 1, Min 1</i>	< 10%	37.26/53	0.007	2.08	0.0414	2.4	95.02%
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<i>Scenario 1, Min 2</i>	< 30%	38.54/53	0.011	2.08	0.000	2.2	93.21%
<i>Scenario 2, Min 1</i>	< 10%	38.54/54	0.006	2.32	0.0113	2.5	93.20%
<i>Scenario 2, Min 2</i>	< 10%	39.05/54	0.004	2.32	0.0003	2.4	93.73%
<i>Scenario 2, Min 1</i>	< 30%	38.33/54	0.035	2.32	0.0023	2.5	94.73%
<i>Scenario 2, Min 2</i>	< 30%	38.80/54	0.025	2.32	0*	2.4	94.09%
<i>Scenario 3</i>	< 10%	39.50/58	0.150	3.65	0.0835	3.7	97.00%
<i>Scenario 4a, Min 1</i>	< 10%	49.93/57	0.079	2.34	0*	1.2	73.52%
<i>Scenario 4a, Min 2</i>	< 10%	49.93/57	0.079	2.34	0*	1.2	73.52%
<i>Scenario 4a, Min 1</i>	< 30%	44.49/57	0.311	2.66	0*	2.4	88.62%
<i>Scenario 4a, Min 2</i>	< 30%	44.49/57	0.311	2.66	0*	2.4	88.62%
<i>Scenario 4b</i>	< 10%	43.56/55	0.054	2.07	0*	1.9	86.70%
<i>Scenario 4b</i>	< 30%	40.03/55	0.218	2.52	0*	2.5	93.54%
<i>Scenario 5a</i>	< 10%	39.39/57	0*	3.22	0.0981	3.2	96.36%
<i>Scenario 5b</i>	< 10%	39.37/55	0*	3.34	0.0060	2.6	94.47%
<i>Scenario 6</i>	< 10%	44.20/58	0*	3.34	0*	2.9	90.93%
<i>Scenario 7a</i>	< 10%	39.21/57	0.126	3.22	0.0616	3.3	96.53%
<i>Scenario 7b</i>	< 10%	39.06/55	0.014	2.56	0.0112	2.7	94.87%
<i>Scenario 8</i>	< 10%	47.32/57	0.259	2.56	0*	1.9	81.60%

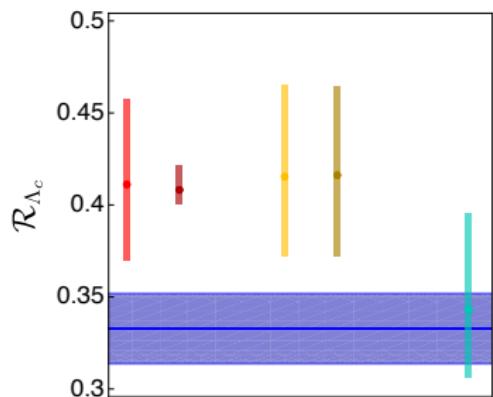
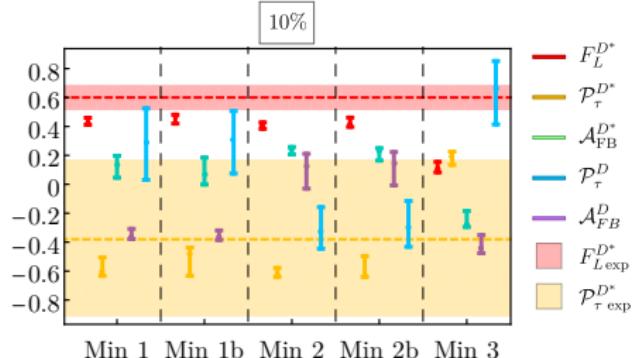
Fit results (RHN)

- Scenarios with several operators: best fits correspond to solutions where all Wilson coefficients but one (i.e. C_{RR}^V) are compatible with zero
- Preference for solutions with all left-handed Wilson coefficients compatible with zero within 1σ
- Preferred solutions: V^μ, S_1, U_1^μ (cannot explain $\bar{F}_L^{D^*}$)
- Only RHN + SM and Φ with $\mathcal{B}(B_c \rightarrow \tau\nu) < 30\%$ can explain $\bar{F}_L^{D^*}$

Predictions for $B \rightarrow D^{(*)}\tau\bar{\nu}_\tau$ observables

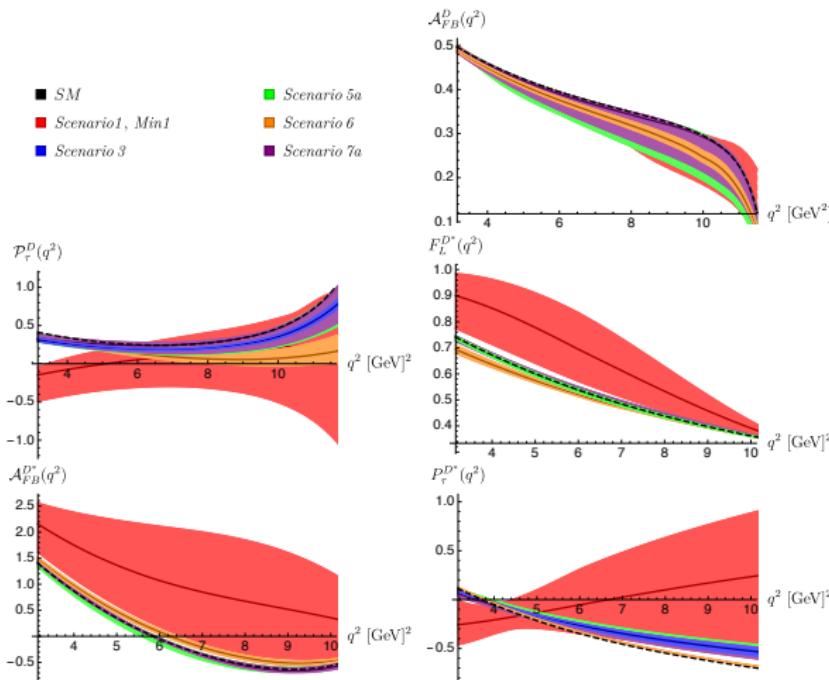


Predictions for $B \rightarrow D^{(*)}\tau\bar{\nu}_\tau$ observables



$$\mathcal{R}_{\Lambda_c} = \frac{\mathcal{B}(\Lambda_b \rightarrow \Lambda_c \tau \bar{\nu}_\tau)}{\mathcal{B}(\Lambda_b \rightarrow \Lambda_c \ell \bar{\nu}_\ell)}.$$

Other predictions

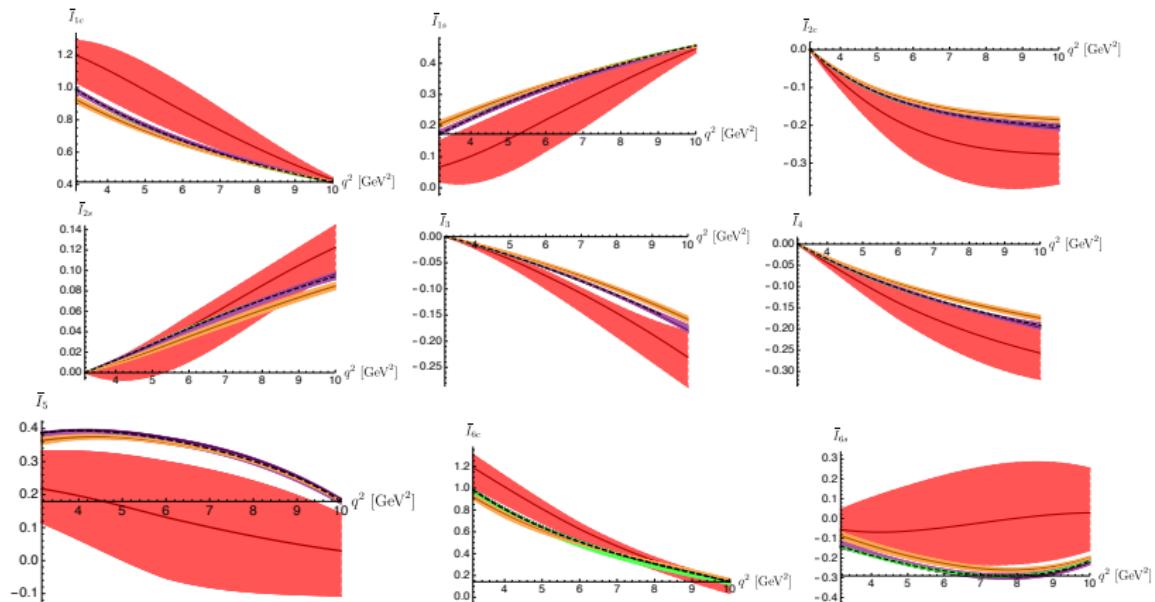


Other predictions

$$\begin{aligned}
 & \frac{d^4\Gamma(B \rightarrow D^*\tau\bar{\nu})}{dq^2 d\cos\theta_\tau d\cos\theta_D d\phi} \equiv I(q^2, \theta_\tau, \theta_D, \phi) \\
 &= \frac{9}{32\pi} \left\{ I_1^s \sin^2 \theta_D + I_1^c \cos^2 \theta_D \right. \\
 &\quad + (I_2^s \sin^2 \theta_D + I_2^c \cos^2 \theta_D) \cos 2\theta_\tau \\
 &\quad + (I_3 \cos 2\phi + I_9 \sin 2\phi) \sin^2 \theta_D \sin^2 \theta_\tau \\
 &\quad + (I_4 \cos \phi + I_8 \sin \phi) \sin 2\theta_D \sin 2\theta_\tau \\
 &\quad + (I_5 \cos \phi + I_7 \sin \phi) \sin 2\theta_D \sin \theta_\tau \\
 &\quad \left. + (I_6^s \sin^2 \theta_D + I_6^c \cos^2 \theta_D) \cos \theta_\tau \right\}
 \end{aligned}$$

$$\bar{I}_i(q^2) \equiv \frac{I_i(q^2)}{\Gamma_f(q^2)}.$$

Other predictions



Conclusions

- Global fit to available data in $b \rightarrow c\tau\bar{\nu}_\tau$ transitions
- EFT approach with minima assumptions
 - NP enters only in 3rd generation of fermions
 - There is a sizeable gap between EW scale and NP
 - Operators are $SU(2)_L \otimes U(1)_Y$ invariant and electroweak symmetry breaking is linearly realized
 - All Wilson coefficients are real
- BaBar and Belle q^2 distributions included. Effect of $F_L^{D^*}$ analyzed
- Different fits performed
 - Main fit (without $F_L^{D^*}$): Three minima, one SM-like and two with stronger deviations from the SM
 - Fit with $F_L^{D^*}$: One minimum disappears , tension at 1σ
 - Fit with C_{RL}^V : The tension disappears for fine-tuned solutions
 - Several fits with RHN

Thank you!

2d plots

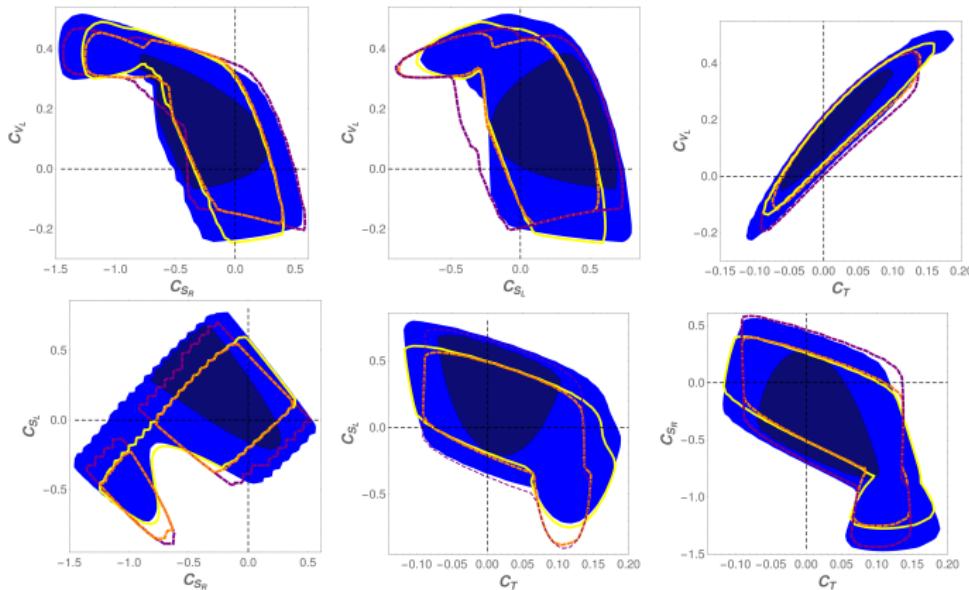


Figure: Blue areas (lighter 95% and darker 68% CL) show the minima without $F_L^{D^*}$ and with $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) \leq 30\%$. The yellow lines display how the 95% CL bounds change when $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) \leq 10\%$. The dashed lines show the effect of adding the observable $F_L^{D^*}$ for both $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) \leq 30\%$ (purple) and for $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) \leq 10\%$ (orange).

Form factors

Parameter	Value
ρ^2	1.32 ± 0.06
c	1.20 ± 0.12
d	-0.84 ± 0.17
$\chi_2(1)$	-0.058 ± 0.020
$\chi'_2(1)$	0.001 ± 0.020
$\chi'_3(1)$	0.036 ± 0.020
$\eta(1)$	0.355 ± 0.040
$\eta'(1)$	-0.03 ± 0.11
$l_1(1)$	0.14 ± 0.23
$l_2(1)$	2.00 ± 0.30

[M. Jung and D. Straub, 2019]

Form Factors

$$\omega(q^2) = \frac{m_B^2 + m_{D^{(*)}}^2 - q^2}{2m_B m_{D^{(*)}}} \quad \text{and} \quad z(q^2) = \frac{\sqrt{\omega(q^2) + 1} - \sqrt{2}}{\sqrt{\omega(q^2) + 1} + \sqrt{2}}.$$

$$\hat{h}_i(q^2) = h_i(q^2)/\xi(q^2)$$

[Bernlochner, 2017]

$$\begin{aligned}\hat{h}_+ &= 1 + \hat{\alpha}_s \left[C_{V_1} + \frac{\omega + 1}{2} (C_{V_2} + C_{V_3}) \right] + (\varepsilon_c + \varepsilon_b) \hat{L}_1, \\ \hat{h}_- &= \hat{\alpha}_s \frac{\omega + 1}{2} (C_{V_2} - C_{V_3}) + (\varepsilon_c - \varepsilon_b) \hat{L}_4, \\ \hat{h}_S &= 1 + \hat{\alpha}_s C_S + (\varepsilon_c + \varepsilon_b) \left(\hat{L}_1 - \hat{L}_4 \frac{\omega - 1}{\omega + 1} \right), \\ \hat{h}_T &= 1 + \hat{\alpha}_s (C_{T_1} - C_{T_2} + C_{T_3}) + (\varepsilon_c + \varepsilon_b) \left(\hat{L}_1 - \hat{L}_4 \right),\end{aligned}$$

Form Factors

$$\begin{aligned}
 \hat{h}_V &= 1 + \hat{\alpha}_s C_{V_1} + \varepsilon_c \left(\hat{L}_2 - \hat{L}_5 \right) + \varepsilon_b \left(\hat{L}_1 - \hat{L}_4 \right), \\
 \hat{h}_{A_1} &= 1 + \hat{\alpha}_s C_{A_1} + \varepsilon_c \left(\hat{L}_2 - \hat{L}_5 \frac{\omega - 1}{\omega + 1} \right) + \varepsilon_b \left(\hat{L}_1 - \hat{L}_4 \frac{\omega - 1}{\omega + 1} \right), \\
 \hat{h}_{A_2} &= \hat{\alpha}_s C_{A_2} + \varepsilon_c \left(\hat{L}_3 + \hat{L}_6 \right), \\
 \hat{h}_{A_3} &= 1 + \hat{\alpha}_s (C_{A_1} + C_{A_3}) + \varepsilon_c \left(\hat{L}_2 - \hat{L}_3 + \hat{L}_6 - \hat{L}_5 \right) + \varepsilon_b \left(\hat{L}_1 - \hat{L}_4 \right), \\
 \hat{h}_P &= 1 + \hat{\alpha}_s C_P + \varepsilon_c \left[\hat{L}_2 + \hat{L}_3 (\omega - 1) + \hat{L}_5 - \hat{L}_6 (\omega + 1) \right] + \varepsilon_b \left(\hat{L}_1 - \right. \\
 \hat{h}_{T_1} &= 1 + \hat{\alpha}_s \left[C_{T_1} + \frac{\omega - 1}{2} (C_{T_2} - C_{T_3}) \right] + \varepsilon_c \hat{L}_2 + \varepsilon_b \hat{L}_1, \\
 \hat{h}_{T_2} &= \hat{\alpha}_s \frac{\omega + 1}{2} (C_{T_2} + C_{T_3}) + \varepsilon_c \hat{L}_5 - \varepsilon_b \hat{L}_4, \\
 \hat{h}_{T_3} &= \hat{\alpha}_s C_{T_2} + \varepsilon_c \left(\hat{L}_6 - \hat{L}_3 \right),
 \end{aligned}$$

Angular observables

$$\frac{d^2\Gamma^{D^{(*)}}}{dq^2 d \cos \theta_\ell} = a_\ell^{(*)}(q^2) - b_\ell^{(*)}(q^2) \cos \theta_\ell + c_\ell^{(*)}(q^2) \cos^2 \theta_\ell ,$$

$$\mathcal{A}_{\text{FB}}^{D^{(*)}}(q^2) \equiv b_\ell^{(*)}(q^2) \left/ \frac{d\Gamma^{D^{(*)}}}{dq^2} \right. = \left(\int_{-1}^0 d \cos \theta_\ell \frac{d^2\Gamma^{D^{(*)}}}{dq^2 d \cos \theta_\ell} - \int_0^1 d \cos \theta_\ell \frac{d^2\Gamma^{D^{(*)}}}{dq^2 d \cos \theta_\ell} \right)$$

$$\mathcal{P}_\tau^{D^{(*)}}(q^2) = \left(\frac{d\Gamma_{\lambda_\tau=1/2}^{D^{(*)}}}{dq^2} - \frac{d\Gamma_{\lambda_\tau=-1/2}^{D^{(*)}}}{dq^2} \right) \left/ \frac{d\Gamma^{D^{(*)}}}{dq^2} \right. ,$$

Angular observables

$$F_L^{D^*}(q^2) = \frac{d\Gamma_{\lambda_{D^*}=0}}{dq^2} \Bigg/ \frac{d\Gamma^{D^*}}{dq^2}.$$

$$\mathcal{O} = \frac{1}{\Gamma^{D^{(*)}}} \int_{m_\tau^2}^{q_{\max}^2} dq^2 \mathcal{O}(q^2),$$

$$\mathcal{P}_\tau^{D^*} = -0.38 \pm 0.51 \text{ (stat)} {}^{+0.21}_{-0.16} \text{ (syst)}$$

$$F_L^{D^*} = 0.60 \pm 0.08 \text{ (stat)} \pm 0.04 \text{ (syst)}$$

Theoretical framework - Observables in the fit

Other fits:

- The ratios $\mathcal{R}_{D^{(*)}}$
- Differential distributions of the decay rates $\Gamma(B \rightarrow D^{(*)} \tau \bar{\nu}_\tau)$
[M. Blanke et. al '18, R. Shi et. al '19, A. Kumar et. al '19 ...]
- The longitudinal polarization fraction $F_L^{D^*}$
- The leptonic decay rate $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}_\tau) \leq 60\%$ [M. Blanke et. al '18]
- $\mathcal{P}_\tau^{D^*}$ and $\mathcal{R}_{J/\psi}$ included
[M. Blanke et. al '18, R. Shi et. al '19, A. Kumar et. al '19 ...]

Fit and results

	Min 1b	Min 2b	Min 1b	Min 2b
$\mathcal{B}(B_c \rightarrow \tau\nu)$	10%		30%	
$\chi^2_{\min}/\text{d.o.f.}$	37.6/54	42.1/54	37.6/54	42.0/54
C_{LL}^V	$0.14^{+0.14}_{-0.12}$	$0.41^{+0.05}_{-0.05}$	$0.14^{+0.14}_{-0.14}$	$0.40^{+0.06}_{-0.07}$
C_{RL}^S	$0.09^{+0.14}_{-0.52}$	$-1.15^{+0.18}_{-0.09}$	$0.09^{+0.33}_{-0.56}$	$-1.34^{+0.57}_{-0.08}$
C_{LL}^S	$-0.09^{+0.52}_{-0.11}$	$-0.34^{+0.13}_{-0.19}$	$-0.09^{+0.68}_{-0.21}$	$-0.18^{+0.13}_{-0.57}$
C_{LL}^T	$0.02^{+0.05}_{-0.05}$	$0.12^{+0.04}_{-0.04}$	$0.02^{+0.05}_{-0.05}$	$0.11^{+0.03}_{-0.04}$

Min 1b compatible with minima found at M. Blanke et. al '18, R. Shi et. al '19, A. Kumar et. al '19 ...

Fit results

Scenario	$\mathcal{B}(B_c \rightarrow \tau \bar{\nu})$	$\chi^2/\text{d.o.f}$	$\bar{\mathcal{P}}_{\tau}^{D^*}, F_L^{D^*}$	Pull _{SM} \mathcal{R}_{D,D^*}	$d\Gamma/dq^2$	Pull _{SM}	p-value
SM	2.16%	52.87/59					69.95%
<i>Scenario 1, Min 1</i>	< 10%	37.26/53	0.007	2.08	0.0414	2.4	95.02%
<i>Scenario 1, Min 2</i>	< 10%	38.86/53	0.001	2.08	0.0006	2.2	92.68%
<i>Scenario 1, Min 1</i>	< 30%	36.42/53	0.022	2.08	0.0866	2.5	96.00%
<i>Scenario 1, Min 2</i>	< 30%	38.54/53	0.011	2.08	0.000	2.2	93.21%
<i>Scenario 2, Min 1</i>	< 10%	38.54/54	0.006	2.32	0.0113	2.5	93.20%
<i>Scenario 2, Min 2</i>	< 10%	39.05/54	0.004	2.32	0.0003	2.4	93.73%
<i>Scenario 2, Min 1</i>	< 30%	38.33/54	0.035	2.32	0.0023	2.5	94.73%
<i>Scenario 2, Min 2</i>	< 30%	38.80/54	0.025	2.32	0*	2.4	94.09%
<i>Scenario 3</i>	< 10%	39.50/58	0.150	3.65	0.0835	3.7	97.00%
<i>Scenario 4a, Min 1</i>	< 10%	49.93/57	0.079	2.34	0*	1.2	73.52%
<i>Scenario 4a, Min 2</i>	< 10%	49.93/57	0.079	2.34	0*	1.2	73.52%
<i>Scenario 4a, Min 1</i>	< 30%	44.49/57	0.311	2.66	0*	2.4	88.62%
<i>Scenario 4a, Min 2</i>	< 30%	44.49/57	0.311	2.66	0*	2.4	88.62%
<i>Scenario 4b</i>	< 10%	43.56/55	0.054	2.07	0*	1.9	86.70%
<i>Scenario 4b</i>	< 30%	40.03/55	0.218	2.52	0*	2.5	93.54%
<i>Scenario 5a</i>	< 10%	39.39/57	0*	3.22	0.0981	3.2	96.36%
<i>Scenario 5b</i>	< 10%	39.37/55	0*	3.34	0.0060	2.6	94.47%
<i>Scenario 6</i>	< 10%	44.20/58	0*	3.34	0*	2.9	90.93%
<i>Scenario 7a</i>	< 10%	39.21/57	0.126	3.22	0.0616	3.3	96.53%
<i>Scenario 7b</i>	< 10%	39.06/55	0.014	2.56	0.0112	2.7	94.87%
<i>Scenario 8</i>	< 10%	47.32/57	0.259	2.56	0*	1.9	81.60%